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Measurement of jet p_T correlations in Pb+Pb and pp collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector

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ABSTRACT

Measurements of dijet p_T correlations in Pb+Pb and pp collisions at a nucleon-nucleon centre-of-mass energy of $\sqrt{s_{NN}} = 2.76$ TeV are presented. The measurements are performed with the ATLAS detector at the Large Hadron Collider using Pb+Pb and pp data samples corresponding to integrated luminosities of 0.14 nb^{-1} and 4.0 pb^{-1} , respectively. Jets are reconstructed using the anti- k_t algorithm with radius parameter values $R = 0.3$ and $R = 0.4$. A background subtraction procedure is applied to correct the jets for the large underlying event present in Pb+Pb collisions. The leading and sub-leading jet transverse momenta are denoted p_{T_1} and p_{T_2} . An unfolding procedure is applied to the two-dimensional (p_{T_1}, p_{T_2}) distributions to account for experimental effects in the measurement of both jets. Distributions of $(1/N)dN/dx_J$, where $x_J = p_{T_2}/p_{T_1}$, are presented as a function of p_{T_1} and collision centrality. The distributions are found to be similar in peripheral Pb+Pb collisions and pp collisions, but highly modified in central Pb+Pb collisions. Similar features are present in both the $R = 0.3$ and $R = 0.4$ results, indicating that the effects of the underlying event are properly accounted for in the measurement. The results are qualitatively consistent with expectations from partonic energy loss models.

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1. Introduction

Jets have long been considered an important tool for studying the matter produced in ultra-relativistic heavy-ion collisions. In these collisions, a hot medium of deconfined colour charges is produced, known as the quark-gluon plasma (QGP). Jets produced in the initial stage of the collision lose energy as they propagate through the medium. This phenomenon, known as *jet quenching*, was first observed at the Relativistic Heavy Ion Collider (RHIC) [1,2]. Early measurements using fully reconstructed jets in Pb+Pb collisions at the LHC provided a direct observation of this phenomenon [3]. In Pb+Pb collisions the transverse momentum (p_T) balance between two jets was found to be distorted, resulting from configurations in which the two jets suffer different amounts of energy loss. This measurement was the experimental confirmation of some of the initial pictures of jet quenching and signatures of a deconfined medium [4].

Subsequent measurements of jets in Pb+Pb collisions have improved the understanding of properties of quenched jets and the empirical features of the quenching mechanism [5–14]. Significant theoretical advances also occurred in this period, and while a complete description of jet quenching is not available, some models are capable of reproducing its key features and provid-

ing testable predictions. Measurements of the dijet asymmetry, $A_J \equiv (p_{T_1} - p_{T_2})/(p_{T_1} + p_{T_2})$, where p_{T_1} and p_{T_2} are the transverse momenta of the jets with the highest and second highest p_T in the event, respectively, have been crucial in facilitating these developments. The experimental results demonstrate that the measured asymmetries in central collisions, where the geometric overlap of the colliding nuclei is almost complete, differ from those in pp collisions more than is expected from detector-specific experimental effects [3,9,10]. However, such effects, in particular the resolution of the measured jet p_T , must be corrected for in order for the measurement to be directly compared to theoretical calculations. Unfolding procedures have been applied to correct for such effects for single-jet measurements [6]; however, the dijet result requires a two-dimensional unfolding to account for migration in the p_T of each jet separately. The measurement reported here is the first unfolded Pb+Pb dijet measurement and as such can be directly compared to theoretical models.

This Letter presents a measurement of dijet p_T correlations in Pb+Pb and pp collisions at a nucleon-nucleon centre-of-mass energy of 2.76 TeV performed with the ATLAS detector. Jets are reconstructed with the anti- k_t algorithm with radius parameter values $R = 0.3$ and $R = 0.4$ [15]. The analysis is described mostly for the example of $R = 0.4$ jets. A background subtraction procedure is applied to account for the effects of the large underlying event (UE) present in Pb+Pb collisions on the measured jet kinematics. The momentum balance of the dijet system is expressed

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by the variable $x_j \equiv p_{T_2}/p_{T_1}$. Measurements of the dijet yield normalised by the total number of jet pairs in a given p_{T_1} interval, $(1/N)dN/dx_j$, are presented as a function of x_j in intervals of p_{T_1} and collision centrality. The results are obtained by first measuring the two-dimensional distribution, (p_{T_1}, p_{T_2}) , and unfolding in the two-dimensional space. The binning in the (p_{T_1}, p_{T_2}) distribution is chosen such that the bins in the two-dimensional space correspond to fixed ranges of x_j , and the $(1/N)dN/dx_j$ results are obtained by projecting into these x_j bins. The (p_{T_1}, p_{T_2}) distributions are less strongly correlated for jets reconstructed with a smaller value of R due to the effects of parton radiation outside the jet cone, which makes them less suitable as a probe of medium-induced effects in Pb+Pb collisions. However, for smaller jet sizes the effect of the UE on the measurement is significantly reduced. It is therefore interesting to compare the results obtained using $R = 0.3$ and $R = 0.4$ jets, to see if the same features are visible.

2. Experimental set-up

The measurements presented in this Letter are performed using the ATLAS inner detector, calorimeter and trigger systems [16]. The inner detector provides measurements of charged-particle tracks over the range $|\eta| < 2.5$.¹ It is composed of silicon pixel detectors in the innermost layers, followed by silicon microstrip detectors and a straw-tube tracker, all immersed in a 2 T axial magnetic field provided by a solenoid. The minimum-bias trigger scintillators (MBTS) measure charged particles over $2.1 < |\eta| < 3.9$ using two planes of counters placed at $z = \pm 3.6$ m and provide timing measurements used in the event selection [17].

The ATLAS calorimeter system consists of a liquid argon (LAr) electromagnetic (EM) calorimeter ($|\eta| < 3.2$), a steel-scintillator sampling hadronic calorimeter ($|\eta| < 1.7$), a LAr hadronic calorimeter ($1.5 < |\eta| < 3.2$), and a forward calorimeter (FCal) ($3.2 < |\eta| < 4.9$). The hadronic calorimeter has three sampling layers longitudinal in shower depth and has a $\Delta\eta \times \Delta\phi$ granularity of 0.1×0.1 for $|\eta| < 2.5$ and 0.2×0.2 for $2.5 < |\eta| < 4.9$.² The EM calorimeters are longitudinally segmented in shower depth into three compartments following a pre-sampler layer ($|\eta| < 1.8$). The EM calorimeter has a granularity that varies with layer and pseudorapidity, but which is generally much finer than that of the hadronic calorimeter. The first layer has high η granularity (between 0.003 and 0.006) that can be used to identify photons and electrons. The middle sampling layer, which typically has the largest energy deposit in EM showers, has a granularity of 0.025×0.025 over $|\eta| < 2.5$. A total transverse energy (TE) trigger is implemented by requiring a hardware-based determination of the total transverse energy in the calorimeter system, E_T^{tot} , to be above a threshold.

The zero-degree calorimeters (ZDCs) are located symmetrically at $z = \pm 140$ m and cover $|\eta| > 8.3$. In Pb+Pb collisions the ZDCs primarily measure “spectator” neutrons: neutrons that do not interact hadronically when the incident nuclei collide. A ZDC coincidence trigger is implemented by requiring the pulse height from each ZDC to be above a threshold set below the single-neutron peak.

In addition to the ZDC and TE hardware-based triggers, a software-based high-level trigger is used to further reduce the ac-

cepted event rate. This trigger applies a jet reconstruction procedure, including a UE subtraction, similar to that used in the offline analysis, which is described in Section 4.

3. Data and Monte Carlo samples

The Pb+Pb data used for these measurements were recorded in 2011 and obtained using a combination of jet and minimum-bias triggers. The minimum-bias trigger is defined by a logical OR of the TE trigger with a threshold of $E_T^{\text{tot}} = 50$ GeV and the ZDC coincidence trigger. The combined trigger is fully efficient in the range of centralities presented here. In the events selected by the ZDC coincidence trigger alone, at least one track is required to remove empty events. The jet trigger [18] first selects events satisfying the TE trigger with a threshold of $E_T^{\text{tot}} = 20$ GeV. A jet reconstruction procedure is then applied using the anti- k_t algorithm with $R = 0.2$ and utilising a UE subtraction procedure similar to that used in the offline reconstruction described in Section 4. Events with at least one jet with $E_T > 20$ GeV at the electromagnetic scale [19] are selected by the jet trigger. The use of $R = 0.2$ for jets in the trigger, as opposed to the values of $R = 0.3$ and 0.4 applied in the measurement, is motivated by the need to define an algorithm that is robust against UE fluctuations, which grow with R . The effects of the different R values on the trigger efficiency are discussed in Section 5. The minimum-bias trigger operated with a prescale of approximately 18 while no prescale was applied to the jet trigger. After accounting for these prescales, the recorded events correspond to integrated luminosities of $8 \mu\text{b}^{-1}$ and 0.14 nb^{-1} for the minimum-bias and jet-triggered samples, respectively.

Events are further subjected to criteria designed to remove non-collision background and inelastic electromagnetic interactions between the nuclei. Events are required to have a reconstructed primary vertex and have a timing difference of less than 5 ns between the times measured by the two MBTS planes. After the trigger and event selection criteria, the resulting data samples contain 53 and 14 million events in the minimum-bias and jet triggered samples, respectively. The average number of collisions per bunch-crossing in the Pb+Pb data sample was less than 0.001, and the effects of multiple collisions are neglected in the data analysis.

The centrality of the Pb+Pb collisions is characterised by the total transverse energy measured in the FCal modules, $\sum E_T^{\text{FCal}}$. The $\sum E_T^{\text{FCal}}$ distribution obtained in minimum-bias collisions is partitioned into separate ranges of $\sum E_T^{\text{FCal}}$ referred to as centrality classes [17,20,21]. Each class is defined by the fraction of the distribution contained by the interval, e.g. the 0–10% centrality class, which corresponds to the most central collisions, contains the 10% of minimum-bias events with the largest $\sum E_T^{\text{FCal}}$. The centrality boundaries used in this analysis are 0%, 10%, 20%, 30%, 40%, 60% and 80%.

The pp data sample, recorded in 2013, was composed of events selected by a jet trigger and used a series of different p_T thresholds each selected with a different prescale. The jet trigger is the same used in other ATLAS measurements in pp collisions [18] and applies the anti- k_t algorithm with $R = 0.4$. The events are further required to contain at least one primary reconstructed vertex. The average number of pp collisions per bunch-crossing varied between 0.3 and 0.6 during data taking. The sample corresponds to a luminosity of 4.0 pb^{-1} .

The impact of experimental effects on the measurement is evaluated using the GEANT4-simulated detector response [22,23] in a Monte Carlo (MC) sample of pp hard-scattering events. Dijet events at $\sqrt{s} = 2.76$ TeV are generated using PYTHIA version 6.423 [24] with parameter values chosen according to the AUET2B tune [25] using the CTEQ6L1 parton distribution function (PDF) set [26]. To

¹ ATLAS uses a right-handed coordinate system with its origin at the nominal interaction point (IP) in the centre of the detector and the z -axis along the beam pipe. The x -axis points from the IP to the centre of the LHC ring, and the y -axis points upward. Cylindrical coordinates (r, ϕ) are used in the transverse plane, ϕ being the azimuthal angle around the beam pipe. The pseudorapidity is defined in terms of the polar angle θ as $\eta = -\ln \tan(\theta/2)$.

² An exception is the third sampling layer, which has a segmentation of 0.2×0.1 up to $|\eta| = 1.7$.

fully populate the kinematic range considered in the measurement, hard-scattering events are generated for separate intervals of \hat{p}_T , the transverse momentum of outgoing partons in the $2 \rightarrow 2$ hard-scattering, and combined using weights proportional to their respective cross sections. Separate samples are generated for the Pb+Pb and pp analyses, with the simulated detector conditions chosen to match those present during the recording of the respective data samples. In the pp data sample, the contribution of additional collisions in the same bunch crossing (pile-up) is accounted for by overlaying minimum-bias pp collisions produced at the same rate as in the data, generated by Pythia version 8.160 [27] using the A2 [28] tune with CT10 PDF set [29]. In the Pb+Pb sample, the UE contribution to the detector signal is accounted for by overlaying the simulated events with minimum-bias Pb+Pb data. The vertex position of each simulated event is selected to match the data event that is overlaid. Through this procedure the MC sample contains contributions from underlying-event fluctuations and harmonic flow that match those present in the data. The combined signal is then reconstructed using the same procedure as is applied to the data. So-called *truth jets* are defined by applying the anti- k_t algorithm with $R = 0.3$ and $R = 0.4$ to stable particles in the MC event generator's output, defined as those with a proper lifetime greater than 10 ps, but excluding muons and neutrinos, which do not leave significant energy deposits in the calorimeter.

The detector's response to quenched jets is studied with an additional sample using PyQuen [30]. This event generator applies medium-induced energy loss to parton showers produced by Pythia. It is used to generate a sample of jets with fragmentation functions that differ from those in the nominal Pythia sample in a fashion consistent with measurements of fragmentation functions in quenched jets [11–13].

4. Jet reconstruction

The procedure used to reconstruct jets in heavy-ion collisions is described in detail in Ref. [5] and is briefly summarised here. First, energy deposits in the calorimeter cells are assembled into $\Delta\eta \times \Delta\phi = 0.1 \times \frac{\pi}{32}$ logical towers. Jets are formed from the towers by applying the anti- k_t algorithm [15] as implemented in the FastJet software package [31].

An estimate of the UE contribution to each tower within the jet is performed on an event-by-event basis by estimating the transverse energy density, $\rho(\eta, \phi)$. Global azimuthal modulation in the UE arises due to the physics of flow and is traditionally described in terms of the Fourier expansion of the ϕ dependence of the transverse energy density. In the subtraction procedure, the UE estimate is assigned a ϕ dependence using the measured magnitudes and phases of the modulation:

$$\rho(\eta, \phi) = \bar{\rho}(\eta) \times \left(1 + 2 \sum_n v_n \cos[n(\phi - \Psi_n)] \right), \quad (1)$$

where v_n and Ψ_n are the magnitudes and phases of the harmonic modulation, respectively, and $\bar{\rho}(\eta)$ is the average transverse energy density measured from energy deposits in the calorimeter as a function of η . In Ref. [5], only the second-order harmonic modulation ($n = 2$) was considered, but in this measurement the procedure has been extended to account for $n = 3$ and 4 harmonic modulations as well. The subtraction is applied to each tower within the jet. The quantities in Eq. (1) may be biased if the energy in a jet is included in their calculation, which results in an over-subtraction of the average UE contribution to the jet energy or incomplete removal of the harmonic modulation. To mitigate such effects, the contribution from jets is excluded from the estimate of the background. The typical background energy subtracted from

the jets varies from a few GeV in peripheral collisions to 150 GeV in the most central collisions.

A calibration factor, derived from MC studies, is then applied after the subtraction to account for the non-compensating hadronic response. A final *in situ* calibration is applied to account for known differences in detector response between data and the MC sample used to derive the initial calibration [32]. This calibration is derived in 8 TeV pp data and adapted to the different beam energy and pile-up conditions relevant for the samples considered here. It uses the balance between jet pairs in different η regions of the detector to provide an evaluation of the relative response to jets as a function of η . It subsequently uses jets recoiling against objects with an independently-determined energy scale such as Z bosons or photons to provide constraints on the absolute energy measurement.

5. Data analysis

In this analysis, jet pairs are formed from the two highest- p_T jets in the event with $p_T > 25$ GeV and $|\eta| < 2.1$. The pair is required to have $\Delta\phi > 7\pi/8$, where $\Delta\phi \equiv |\phi_1 - \phi_2|$. For events selected by a jet trigger, the leading jet is required to match a jet identified by the trigger algorithm responsible for selecting the jet. The two-dimensional (p_{T_1}, p_{T_2}) distributions obtained from different triggered samples are combined such that intervals of p_{T_1} are populated by a single trigger. In the pp data analysis, the trigger with the most events that is more than 99% efficient for selecting a jet with $p_T > p_{T_1}$ is used, with the reciprocal of the luminosity for the respective trigger samples used as a weight.

The Pb+Pb jet trigger efficiency has a broad turn-on as a function of p_T since the trigger jets are identified using $R = 0.2$ and have no energy scale calibration applied. This effect is the strongest in central collisions where the UE fluctuations are the largest and further weaken the correlation between jets reconstructed with different values of R . In the most central collisions, the single-jet-trigger efficiency does not reach a plateau until $p_T \sim 90$ GeV. The jet-triggered sample is used where the efficiency is found to be greater than 97%, which occurs at a p_T of approximately 85 GeV in the most central collisions. A trigger efficiency correction is applied in the region where there is an inefficiency.

In addition to the dijet signal, the measured (p_{T_1}, p_{T_2}) distribution receives contributions from so-called *combinatorial* jet pairs. Such pairs arise when two jets, which are not from the same hard-scattering process, fulfil the pair requirements through random association. Jets forming such pairs may originate from independent hard scatterings or from upward UE fluctuations identified as jets, referred to as UE jets. The rate for such occurrences is highest in the most central collisions, and the reduction in the true sub-leading jet p_T due to quenching effects further enhances the likelihood of forming a combinatoric pair.

The shape of the $\Delta\phi$ distribution for the combinatoric jet pairs is influenced by the harmonic flow. Since the jet p_T spectrum falls steeply, the jets most likely to be measured at a given p_T value are those lying on top of larger-than-average UE. If the effects of the modulation of the UE are not fully accounted for in the background subtraction, more jets would be observed at angles corresponding to the flow maxima ($\phi \sim \Psi_n$). Thus combinatoric jet pairs, without any underlying angular correlation, are expected to acquire a modulation to their $\Delta\phi$ distribution determined by the dominant flow harmonics [33]. Although the second-, third- and fourth-order harmonic modulations are considered event-by-event in the jet reconstruction procedure described in Section 4, only the effects of the second-order modulation on the $\Delta\phi$ distribution are observed to be completely removed. The residual effects are an indication that the method of estimating the modulation of the UE under-

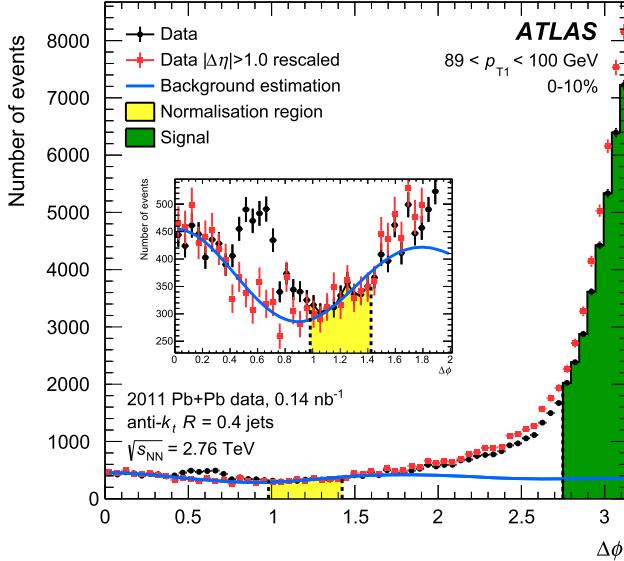


Fig. 1. The $\Delta\phi$ distribution for $R = 0.4$ jet pairs with $89 < p_{T_1} < 100$ GeV in the 0–10% centrality interval. The distribution for all jet pairs is indicated by black circles. The combinatoric contribution given by Eq. (2) is shown as a blue line. The ranges of $\Delta\phi$ used to fix the value of Y and to define the signal region ($\Delta\phi > \frac{7\pi}{8}$) are indicated by yellow and green shaded regions, respectively. The parameters c_3 and c_4 are obtained by fitting the $\Delta\phi$ distribution for jet pairs with $|\Delta\eta| > 1$ in the region $0 < \Delta\phi < \frac{\pi}{2}$, which is indicated by red squares (scaled to match the black circles in the yellow region for presentation purposes). The error bars denote statistical errors. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

neath the jet is less accurate for the higher-order harmonics than for $n = 2$.

To account for the residual modulation, the combinatoric contribution is assumed to be of the form:

$$C(\Delta\phi) = Y(1 + 2c_3 \cos 3\Delta\phi + 2c_4 \cos 4\Delta\phi). \quad (2)$$

The c_3 and c_4 values are determined by fitting the $\Delta\phi$ distributions over the range $0 < \Delta\phi < \pi/2$ where the real dijet contribution is expected to be small. The region $0 < \Delta\phi \lesssim 0.8$ is also expected to receive real dijet contributions arising from parton radiation which results in pairs of jets at nearby angles. To remove this contribution, the fit to obtain c_3 and c_4 is performed only using jet pairs with a separation of $|\Delta\eta| > 1$. Once c_3 and c_4 are obtained, the $\Delta\phi$ distribution without this $|\Delta\eta|$ requirement is integrated over the range $1 < \Delta\phi < 1.4$ to obtain Y . This proce-

dure is performed separately in each (p_{T_1}, p_{T_2}) interval. In intervals where the c_3 and c_4 are found to not be statistically significant their values are taken to be zero. The expected combinatorial contribution, B , in the signal region is obtained by integrating $C(\Delta\phi)$ from $7\pi/8$ to π .

The $\Delta\phi$ distribution of jet pairs is shown in Fig. 1 for pairs with $89 < p_{T_1} < 100$ GeV in the 0–10% centrality interval. Also shown is the $\Delta\phi$ distribution obtained from such jet pairs with $|\Delta\eta| > 1$, which is fitted to obtain c_3 and c_4 . The background subtraction is most significant in central collisions, where the fraction subtracted from the total yield in the signal region is as large as 10% for small x_j and is less than 1% for x_j values greater than 0.5. The background contribution in more peripheral collisions is less than 1% for all values of x_j . This background subtraction is not applied in the pp data because the pile-up is small.

The presence of combinatoric jet pairs also reduces the efficiency for genuine pairs. The measured inclusive jet spectrum is used to estimate the likelihood that another jet in the event, uncorrelated with the dijet system, is measured with a transverse momentum greater than p_{T_2} . For the 40–60% and 60–80% centrality intervals the effect is negligible. In the 0–10% centrality bin the efficiency is approximately 0.9 for $p_{T_2} = 25$ GeV and increases with p_{T_2} , reaching unity at 45 GeV. The effects of the combinatoric jet pairs are accounted for by first subtracting the estimated background and then correcting for the efficiency, ε , in each (p_{T_1}, p_{T_2}) bin. The number of jet pairs corrected for such effects is defined to be:

$$N^{\text{corr}} = \frac{1}{\varepsilon} (N^{\text{raw}} - B),$$

where N^{raw} is the number of jet pairs after correcting for trigger efficiency and luminosity/prescale weighting as described above.

In a given event, the p_T resolution may result in the jet with the highest true p_T being measured with the second highest p_T and vice-versa. To properly account for such migration effects, (p_{T_1}, p_{T_2}) distributions are symmetrised prior to the unfolding by apportioning half of the yield in a given (p_{T_1}, p_{T_2}) bin, after combinatoric subtraction, to the bin related to the original by $p_{T_1} \leftrightarrow p_{T_2}$. The two-dimensional distributions after symmetrisation are shown in Fig. 2 for central and peripheral Pb+Pb collisions and for pp collisions. The choice of binning in (p_{T_1}, p_{T_2}) is motivated by the mapping to the x_j variable, and is described in more detail in the following section.

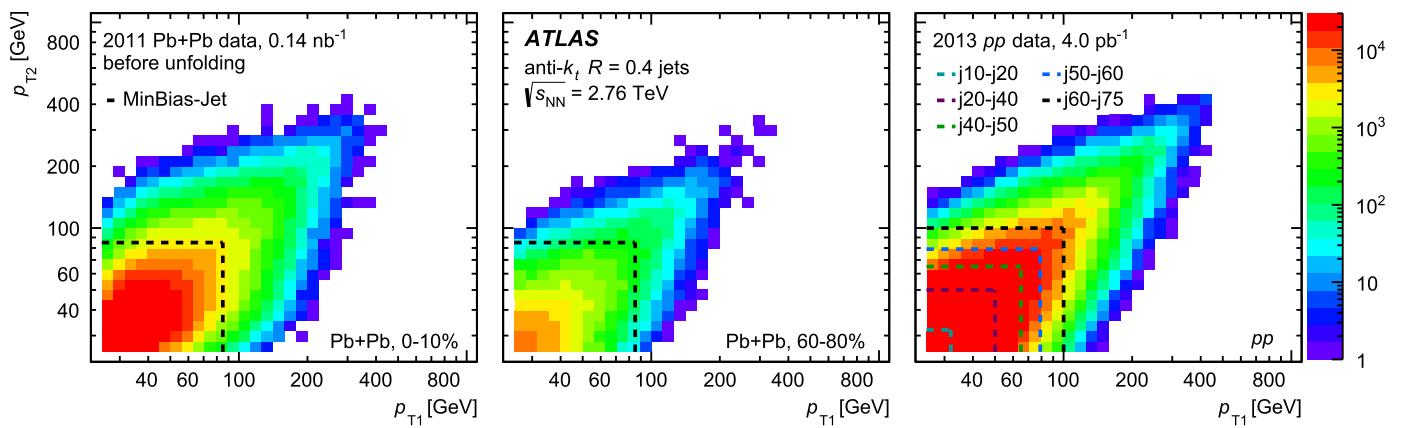


Fig. 2. The two-dimensional (p_{T_1}, p_{T_2}) distributions after correction and symmetrisation for Pb+Pb data in the 0–10% (left) and 60–80% (centre) centrality bins and for pp data (right) for $R = 0.4$ jets. The dashed lines indicate the boundaries used in selecting the different triggers. The Pb+Pb data distributions have their combinatoric contribution subtracted.

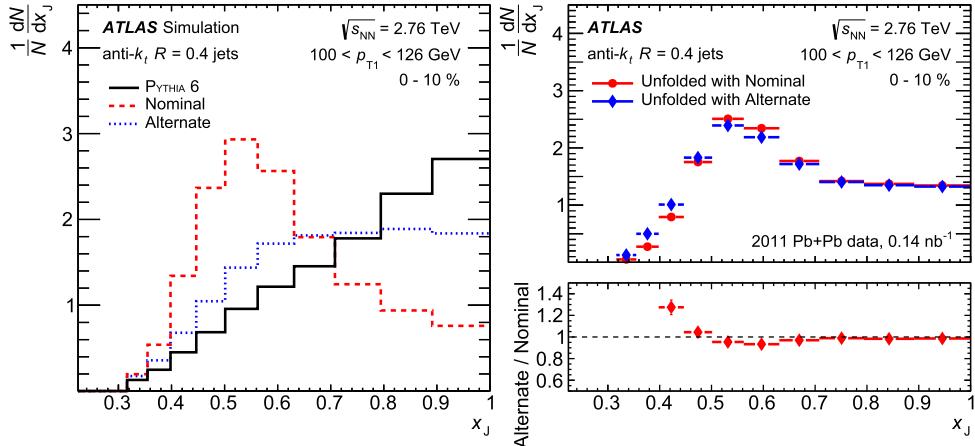


Fig. 3. Left: the $(1/N)dN/dx_J$ distributions used as priors in the unfolding of the $R = 0.4$ jets for the nominal (dashed red) and alternate variation (dotted blue) for the $100 < p_{T_1} < 126$ GeV and 0–10% centrality interval. The same distribution obtained from the PYTHIA MC sample is shown in solid black. Right: unfolded $(1/N)dN/dx_J$ distributions from data for the same p_{T_1} and centrality ranges using the nominal (red circles) and alternate (blue diamonds) priors shown in the left panel. The ratio of nominal to alternate is shown in the bottom panel. In the bottom panel on the right the first two bins are off scale with bins centres of $x_J = 0.34$ and 0.38 and bins contents of 2.49 and 1.82 , respectively. Statistical errors are not shown. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

6. Unfolding

The calorimetric response to jets is evaluated in the MC sample by matching truth and reconstructed jets; the nearest reconstructed and truth jets within $\Delta R = \sqrt{(\Delta\eta)^2 + (\Delta\phi)^2}$ of 0.3 are considered to be a match. The same requirement is applied in both the $R = 0.3$ and $R = 0.4$ versions of the analysis. The response is typically characterised in terms of the jet energy scale (JES) and jet energy resolution (JER). These quantities describe the mean and width of the p_T^{reco} distributions at fixed p_T^{truth} , expressed as a fraction of p_T^{truth} . Generally, the mean of p_T^{reco} differs from p_T^{truth} by less than a percent, independent of p_T^{truth} and centrality. This indicates that the subtraction of the average UE contribution to the jet energy is under good experimental control. The JER receives contributions both from the response of the calorimeter and from local UE fluctuations about the mean in the region of the jet. The latter contribution dominates at low p_T with the resolution as large as 40% at $p_T \simeq 30$ GeV in the most central collisions. At the same p_T , the JER is only 20% in peripheral collisions, similar to that in pp collisions. At larger p_T values the relative contribution of the UE fluctuations to the jet p_T diminishes, and the JER is dominated by detector effects, reaching a constant, centrality-independent value of 8% for $p_T > 300$ GeV.

The migration in the two-dimensional (p_{T_1}, p_{T_2}) distribution is accounted for by applying a two-dimensional Bayesian unfolding to the data [34,35]. This procedure utilizes a response matrix obtained by applying the same pair selections to the truth jets in MC simulation as in the data analysis (except the trigger requirement) and recording the values of $p_{T_1}^{\text{truth}}$ and $p_{T_2}^{\text{truth}}$ and the transverse momenta of the corresponding reconstructed jets $p_{T_1}^{\text{reco}}$ and $p_{T_2}^{\text{reco}}$. The matched reconstructed jets are not required to have the highest p_T in the event, but are subject to all other requirements applied to the data and truth jets. The response matrix is populated symmetrically in both truth and reconstructed p_T . The full four-dimensional response behaves similarly to the factorised product of separate single-jet response distributions, and the migration effects can be understood in terms of the above discussion. While this provides intuition for the nature of the unfolding problem, such a factorisation is not explicitly assumed, and any correlations between the response of the two jets are accounted for in the procedure.

After unfolding, the leading/sub-leading distinction is restored by reflecting the distribution over the line $p_{T_1} = p_{T_2}$: for each bin with $p_{T_2} > p_{T_1}$ the yield is moved to the corresponding bin with $p_{T_2} < p_{T_1}$. The bins along the diagonal, e.g. those containing pairs with $p_{T_2} = p_{T_1}$, are not affected by this procedure. The two-dimensional distribution is constructed using binning along each axis such that the upper edge of the i th bin obeys,

$$p_{T,i} = p_{T,0} \alpha^i, \quad \alpha = \left(\frac{p_{T,N}}{p_{T,0}} \right)^{1/N},$$

where N is the total number of bins and $p_{T,0}$ and $p_{T,N}$ are the minimum and maximum bin edges covered by the binning, respectively. As a consequence, the bins are of the same size when plotted with logarithmic axes. With these choices of binning, the range of x_J values in any given (p_{T_1}, p_{T_2}) bin is fully contained within two adjacent x_J bins, which have boundaries at $x_{J,i} = \alpha^{i-N}$. In this analysis, half of the yield in each (p_{T_1}, p_{T_2}) bin is apportioned to each of the x_J bins. The exceptions are the bins along the diagonal. These bins contribute solely to the x_J bin with bin edges $(\alpha^{-1}, 1)$. The effects of such a mapping on the x_J distribution are studied and found to not significantly distort the shape of the distribution for a variety of input x_J distributions.

The Bayesian unfolding method is an iterative procedure that requires both a choice in a number of iterations, n_{iter} , and assumption of a prior for the underlying true distribution. An increase in n_{iter} reduces sensitivity to the choice of prior but may amplify statistical fluctuations that are already present in the input distribution. As PYTHIA does not include the effects of jet quenching, the x_J distributions obtained from the MC sample are not expected to be optimal choices for the prior. In particular, the x_J distributions in PYTHIA increase monotonically with x_J , whereas the distributions in the data become flatter and develop a peak near $x_J \sim 0.5$ in lower p_{T_1} intervals and in the most central collisions. The (p_{T_1}, p_{T_2}) distributions from PYTHIA are reweighted in a centrality-dependent way to obtain features that qualitatively match those present in the data.

The effects of the reweighting procedure are shown in the left panel of Fig. 3 in the $100 < p_{T_1} < 126$ GeV range and 0–10% centrality interval, where the largest difference between the data and PYTHIA is observed. The “nominal” distribution, or the reweighted distribution, is used as the prior in the unfolding of the data. An

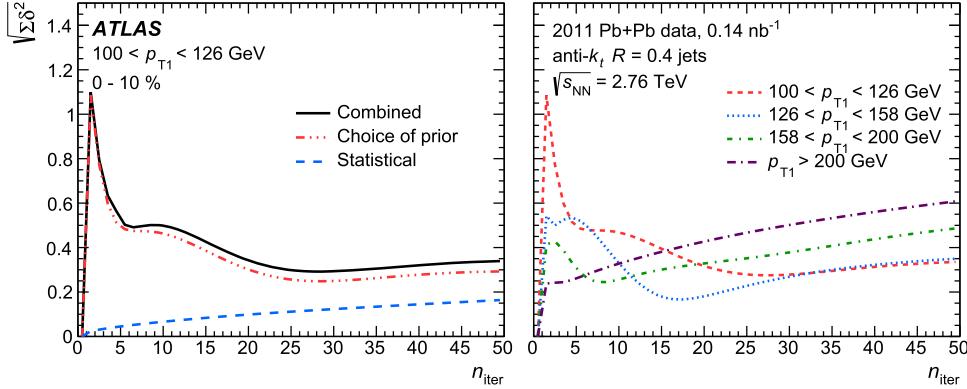


Fig. 4. Uncertainties sensitive to the number of iterations in the unfolding procedure as a function of n_{iter} for the 0–10% centrality interval for $R = 0.4$ jets. Left: The combination (solid black) of the unfolding (dashed red) and statistical (dotted blue) uncertainty, $\sqrt{\Sigma \delta^2}$ for the $100 < p_{T_1} < 126$ GeV interval. Right: The combined uncertainty for each p_{T_1} interval considered in the measurement. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

“alternate” reweighting is also shown, which has a shape significantly different from the nominal, but does not increase as much as the PYTHIA distribution. The features in the data are observed to be robust with respect to the choice of prior for a broad set of reweighting functions. The systematic uncertainty due to the choice of prior is estimated by comparing the results of the unfoldings using the “nominal” and “alternate” x_j distributions. The results of applying unfoldings with these two choices of priors are shown in the right panels of Fig. 3 for the same p_{T_1} and centrality selection.

An alternative study is performed in the MC sample to validate the estimation of this uncertainty. The “alternate” reweighting is applied to obtain input truth and reconstructed distributions in which no peak structure is present. The reconstructed distribution is then unfolded using the *nominal* prior. The unfolded distribution does not develop the strong peak present in the nominal prior. The differences between the unfolded result and the input truth distribution are similar to the uncertainty obtained by varying the prior used to unfold the data.

The value of n_{iter} is selected separately in each centrality interval by examining the uncertainty, $\sqrt{\Sigma \delta^2}$, in $(1/N)dN/dx_j$ after unfolding considering statistical uncertainties and systematic uncertainties attributed to the unfolding procedure,

$$\delta^2 = \delta_{\text{stat}}^2 + \delta_{\text{prior}}^2,$$

and summing over all x_j bins. Here δ_{prior} is the uncertainty due to the choice of prior, obtained using the procedure described above. The statistical uncertainties are evaluated using a pseudo-experiment technique. Stochastic variations of the data are generated based on its statistical uncertainty and each variation is unfolded and projected into x_j . The statistical covariance of the set is taken as the statistical uncertainty. An additional covariance is obtained from applying the pseudo-experiment procedure to the response matrix and combined with that obtained from applying the procedure to the data. The δ_{stat}^2 for each x_j bin is taken to be the diagonal element of the resulting covariance matrix. The statistical covariance matrices exhibit similar trends across all p_{T_1} and centrality ranges. Nearby x_j bins show a strong positive correlation that diminishes for bins separated in x_j , and is expected from the effects of the procedures for unfolding and mapping to x_j . Bins well separated in x_j show an anti-correlation attributable to the normalisation of $(1/N)dN/dx_j$.

The left panel of Fig. 4 shows $\sqrt{\Sigma \delta^2}$ as a function of n_{iter} along with its various contributions for the $100 < p_{T_1} < 126$ GeV range and 0–10% centrality interval. Since the unfolding is performed in two dimensions, the value of n_{iter} cannot be chosen separately for

each range of p_{T_1} . At higher values of p_{T_1} the effects of the unfolding are smaller while the effects of the statistical fluctuations can be more severe. The right panel of Fig. 4 shows the total $\sqrt{\Sigma \delta^2}$ for each range of p_{T_1} considered in the measurement along with the total combined over all p_{T_1} ranges. The value of n_{iter} for each centrality bin and R value is chosen by considering the n_{iter} dependence of $\sqrt{\Sigma \delta^2}$ for each p_{T_1} bin and selecting a value that maintains comparable uncertainties across all p_{T_1} ranges. The more central bins require the most iterations, resulting from the larger jet energy resolution in these events. The number of iterations for $R = 0.4$ jets is at most 20 for 0–10% centrality and at the least 6 for 60–80% centrality. The $\sqrt{\Sigma \delta^2}$ distributions for $R = 0.3$ jets show behaviour similar to those for $R = 0.4$ jets in the same centrality bin.

It is possible for a third jet present in the event to be reconstructed as the jet with the second highest p_T through the experimental resolution. As a check to study the impact of such effects on the measurement, an alternative response matrix is constructed where no ΔR matching is required between the truth and reconstructed jets. A weighting is applied such that the p_T distribution of the reconstructed third jet matches that observed in the data. Differences between the unfolded distributions obtained with this response matrix and the nominal one are observed to be small and well within the systematic uncertainty associated with the unfolding procedure.

The $(1/N)dN/dx_j$ distributions before and after unfolding are shown in Fig. 5 for central and peripheral Pb+Pb collisions and for pp collisions for jet pairs with $100 < p_{T_1} < 126$ GeV. The systematic uncertainties indicated contain all of the contributions to the total systematic uncertainty described in Section 7. In the pp and 60–80% centrality interval, the resolution effects before unfolding reduce the sharpness of the peak near $x_j \sim 1$. In the case of the 0–10% centrality interval, the effect is to smear out the peak near $x_j \sim 0.5$. The lowest x_j bins exhibit instability in the unfolding procedure due to the MC sample having too few events in this region. However, including this range in the unfolding improves the stability of the adjacent x_j bins. Thus, after unfolding, only the range $0.32 < x_j < 1$ is reported in the results even though pairs with $p_{T_2} > 25$ GeV are included in the measurement.

7. Systematic uncertainties

Systematic uncertainties attributed to the response matrix used in the unfolding arise due to uncertainties in the JES and JER. To account for these effects, new response matrices are constructed with a systematically varied relationship between the truth and

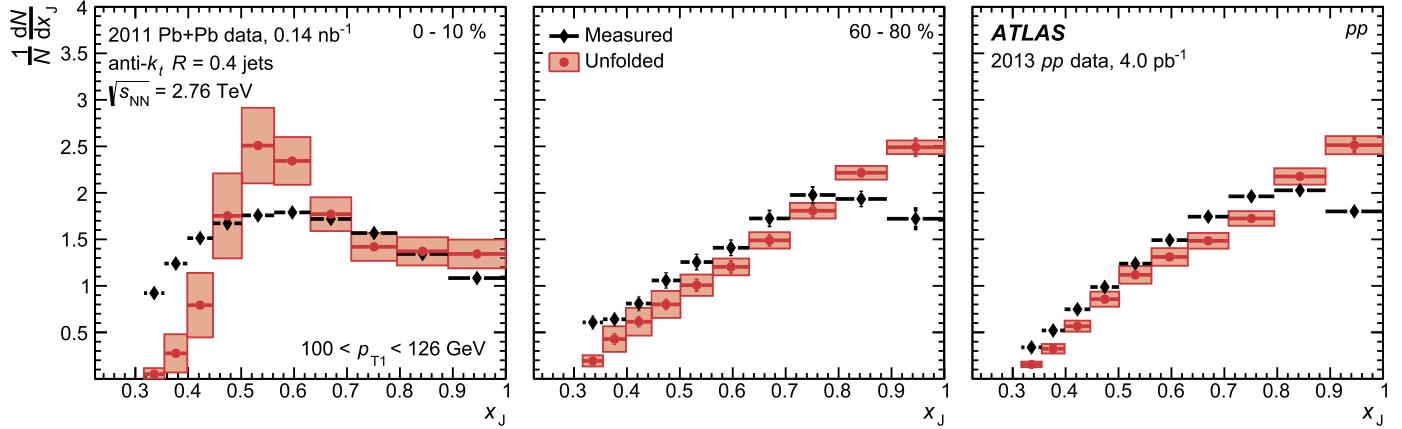


Fig. 5. The $(1/N)dN/dx_J$ distributions for $R = 0.4$ jets before (black) and after (red) unfolding for the $100 < p_{T_1} < 126$ GeV interval for the Pb+Pb 0–10% (left) and Pb+Pb 60–80% (middle) centrality ranges and for pp collisions (right). Statistical uncertainties are indicated by vertical error bars (not visible in most cases). Systematic uncertainties in the unfolded result are indicated by the red shaded boxes. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

reconstructed jet kinematics. The data are then unfolded using the new response and the result is compared with the nominal.

In the pp data analysis, the JES uncertainty is described by a set of 11 independent nuisance parameters; these include effects from uncertainties derived through the *in situ* calibration [32]. In the MC sample used to determine the calibration, the calorimetric response to jets initiated by the fragmentation of quarks and gluons is observed to differ. Potential inaccuracies in the MC sample describing both this flavour-dependent response and the relative abundances of quark and gluon jets are accounted for using separate nuisance parameters. A source of uncertainty related to the adaptation of the *in situ* calibration derived at $\sqrt{s} = 8$ TeV to 2.76 TeV data is also included.

In the Pb+Pb data analysis, two additional uncertainties in the JES are considered. The first accounts for differences between the detector operating conditions in the Pb+Pb and pp data, which were recorded in 2011 and 2013, respectively. This is derived by using charged-particle tracks reconstructed in the inner detector to provide an independent check on the JES, which only uses information from the calorimeter. For each jet, all reconstructed tracks within $\Delta R < 0.4$ and having $p_T^{trk} > 2$ GeV, are matched to the jet and the scalar sum of the track transverse momenta is evaluated. The ratio of this sum to the jet's p_T is evaluated both in data and in the MC sample, and a double ratio of the two quantities is formed. The double ratio obtained in peripheral Pb+Pb data is compared with that in pp data. The precision of the comparison is limited by having too few events in the peripheral Pb+Pb data and at high jet p_T , and a p_T - and η -independent uncertainty of 1.46% is assigned to account for potential differences.

The second additional uncertainty is a centrality-dependent JES uncertainty to account for potential differences in the detector response to quenched jets. This is estimated by comparing the detector response evaluated in the PYTHIA and PyQuen MC samples. This estimate is checked in data using a track-based study similar to the one described above, but comparing central and peripheral Pb+Pb collisions and accounting for the measured variation of the fragmentation function with centrality [11–13]. An uncertainty of up to 1% in the most central collisions and decreasing linearly with centrality percentile to 0% in the 60–80% centrality class is assigned.

The uncertainty attributed to the JER is obtained by adding Gaussian fluctuations to each reconstructed jet p_T value when populating the response matrix. The magnitude of this uncertainty is fixed by a comparison of the data and MC descriptions of the JER

in 8 TeV data [36]. Since the MC sample is constructed using the data overlay procedure, it is expected that the centrality dependence of the JER should be well described in the MC sample. This is checked by studying the distribution of UE fluctuations using random, jet-sized groups of calorimeter towers in Pb+Pb data. The standard deviations of these distributions describe the typical UE contribution beneath a jet. The centrality dependence of the UE fluctuations is compared to that of the JER in the MC sample, and a systematic uncertainty is included to account for the observed differences. As expected, these differences are much smaller than the centrality-independent contribution to the JER uncertainty.

The data-driven estimates of the JES and JER uncertainties described above are derived using $R = 0.4$ jets. Additional uncertainties are included in the $R = 0.3$ jet measurement to account for potential differences between data and the MC sample in the relative energy scale of $R = 0.3$ jets with respect to $R = 0.4$ jets. These uncertainties are estimated from a study that matched jets reconstructed with the two R values and compared the means of the $p_T^{R=0.3}/p_T^{R=0.4}$ distributions in data and the MC sample. Differences may arise between the data and MC sample from differences in the calorimetric response or because the jets in the two samples have different internal structure. The contribution of the latter is constrained by using existing jet shape measurements [37]. An uncertainty in the energy scale is applied to account for residual differences, which are 1.5% at the lowest p_T and decrease sharply as a function of p_T to a limiting value of 0.3% at high p_T . A similar study comparing the variances of the $p_T^{R=0.3}/p_T^{R=0.4}$ distributions is used to constrain the uncertainty in the relative resolution. This uncertainty is applied in the $R = 0.3$ jet measurement in the same fashion as the other JER uncertainties described above. Although larger than the centrality-dependent contribution, it is also much smaller than the centrality-independent contribution.

As the response matrix is sparsely populated (containing 40^4 bins), statistical fluctuations could introduce instabilities in the unfolding. To evaluate the sensitivity to such effects, along with any other defects in the response, a new response matrix is constructed as a factorised product of single-jet response distributions, i.e. assuming the responses in p_{T_1} and p_{T_2} are independent. The data are unfolded using this new response and the differences between the unfolded distributions are taken as a systematic uncertainty. Systematic uncertainties in the unfolding due to the choice of prior are estimated as described in the previous section and are also included.

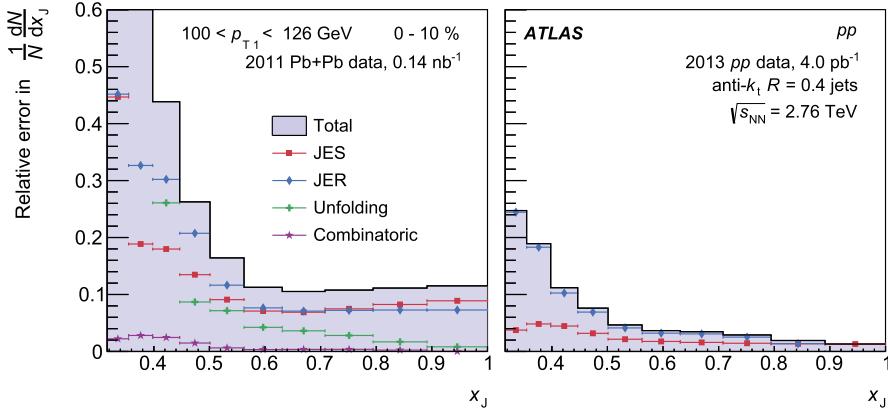


Fig. 6. The total systematic uncertainty and its various components for $100 < p_{T1} < 126$ GeV for $R = 0.4$ jets in Pb+Pb collisions with 0–10% centrality (left) and pp collisions (right). In the figure on the left the first two bins are off scale with bins centres of $x_J = 0.34$ and 0.38 and bins contents of 1.25 and 0.75 , respectively.

Uncertainties due to the correction for the combinatoric effects described in Section 5 affect the number of jet pairs before the unfolding and are thus included as additional contributions to the previously described statistical uncertainties in the data. These include statistical uncertainties in ε and the uncertainties in the values of the fit parameters c_3 and c_4 , accounting for their covariance. Uncertainties in the normalisation are estimated by varying the region of $\Delta\phi$ used to estimate Y from 1.0–1.4 to 1.1–1.5. The uncertainty due to this correction is smaller than the other uncertainties in all p_T and centrality bins, and is only greater than 5% at values of $x_J < 0.4$. This correction was not applied to the pp data so there is no corresponding systematic uncertainty.

The breakdown of different contributions to the total systematic uncertainty is shown in the $100 < p_{T1} < 126$ GeV range for the 0–10% centrality interval and for pp collisions in Fig. 6. Each contribution to the uncertainty, and thus the total uncertainty, tends to decrease with increasing x_J . The total uncertainty at $x_J \sim 1$ reaches approximately 12% in most p_{T1} and centrality bins in the Pb+Pb data. For $x_J < 0.4$, the relative uncertainty becomes large, but this region represents only a small contribution to the total $(1/N)dN/dx_J$ distribution. The JER uncertainty is the largest contribution. In the Pb+Pb data it reaches values of approximately 10% and 15% at $x_J \sim 1$ and $x_J = 0.5$, respectively. The JES contributions are the second largest contribution to the uncertainties, typically between 5% and 10%. In the most central bins the unfolding uncertainty can become as large as the JES contribution. The contributions to the uncertainty in the other centrality intervals and in the pp data follow trends similar to those described for the 0–10% centrality interval, but the magnitudes are smaller in more peripheral collisions. In the pp data they are typically smaller by a factor of two compared to the 0–10% Pb+Pb data. The uncertainties for the $R = 0.3$ result follow the same trends as those for the $R = 0.4$ result but are slightly larger due to the two additional sources included in that measurement to describe the relative energy scale and resolution between the two R values.

8. Results

The unfolded $(1/N)dN/dx_J$ distribution in pp collisions for $100 < p_{T1} < 126$ GeV is shown in Fig. 7. Also shown are the corresponding distributions obtained from the PYTHIA 6 sample used in the MC studies and also from Pythia 8 using the AU2 tune and Herwig++ [38] with the UE-EE-3 [39] tune. An additional sample, referred to as Powheg+Pythia 8 is generated using Powheg-Box 2.0 [40–42], which is accurate to next-to-leading order in perturbative QCD, and interfaced with Pythia 8 to provide a description of the parton shower and hadronisation. All samples used the

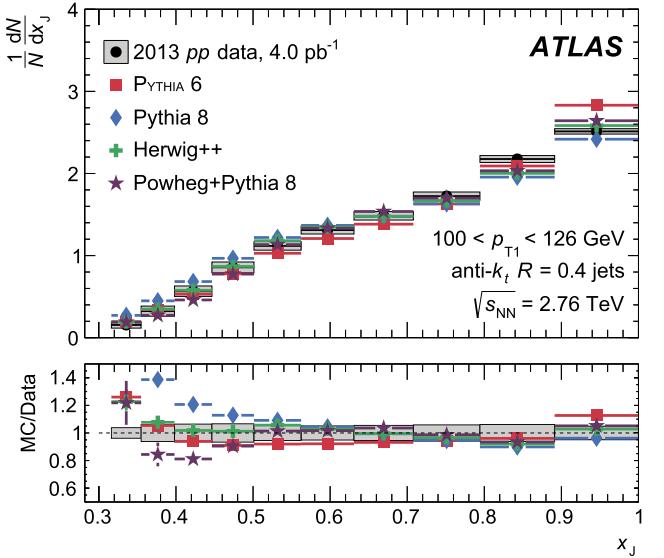


Fig. 7. The $(1/N)dN/dx_J$ distribution for $R = 0.4$ jets in pp collisions for the $100 < p_{T1} < 126$ GeV interval is shown in black points with the grey shaded boxes indicating the systematic uncertainties. Also shown are results obtained from various MC event generators: PYTHIA 6 (red squares), Pythia 8 (blue diamonds), Herwig++ (green crosses) and Powheg+Pythia 8 (purple stars). The ratio of each MC result to the data is shown in the bottom panel where the systematic uncertainties of the data are indicated by a shaded band centred at unity. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

CTEQ6L1 PDF set [26] except the Powheg+Pythia 8, which used the CT10 PDF set [29]. All four models describe the data fairly well with the Herwig++ and Powheg+Pythia 8 showing the best agreement over the full x_J range.

The unfolded $(1/N)dN/dx_J$ distributions in Pb+Pb collisions are shown in Fig. 8, for jet pairs with $100 < p_{T1} < 126$ GeV for different centrality intervals. The distribution in pp collisions is shown on each panel for comparison. In the 60–80% centrality bin, where the effects of quenching are expected to be the smallest, the Pb+Pb data are consistent with the pp data. In more central Pb+Pb collisions, the distributions become significantly broader than that in pp collisions and the peak at $x_J \sim 1$, corresponding to nearly symmetric dijet events, is reduced. At lower centrality percentiles the distribution becomes almost constant over the range $0.6 \lesssim x_J \lesssim 1$, and develops a peak at $x_J \sim 0.5$ in the 0–10% centrality interval.

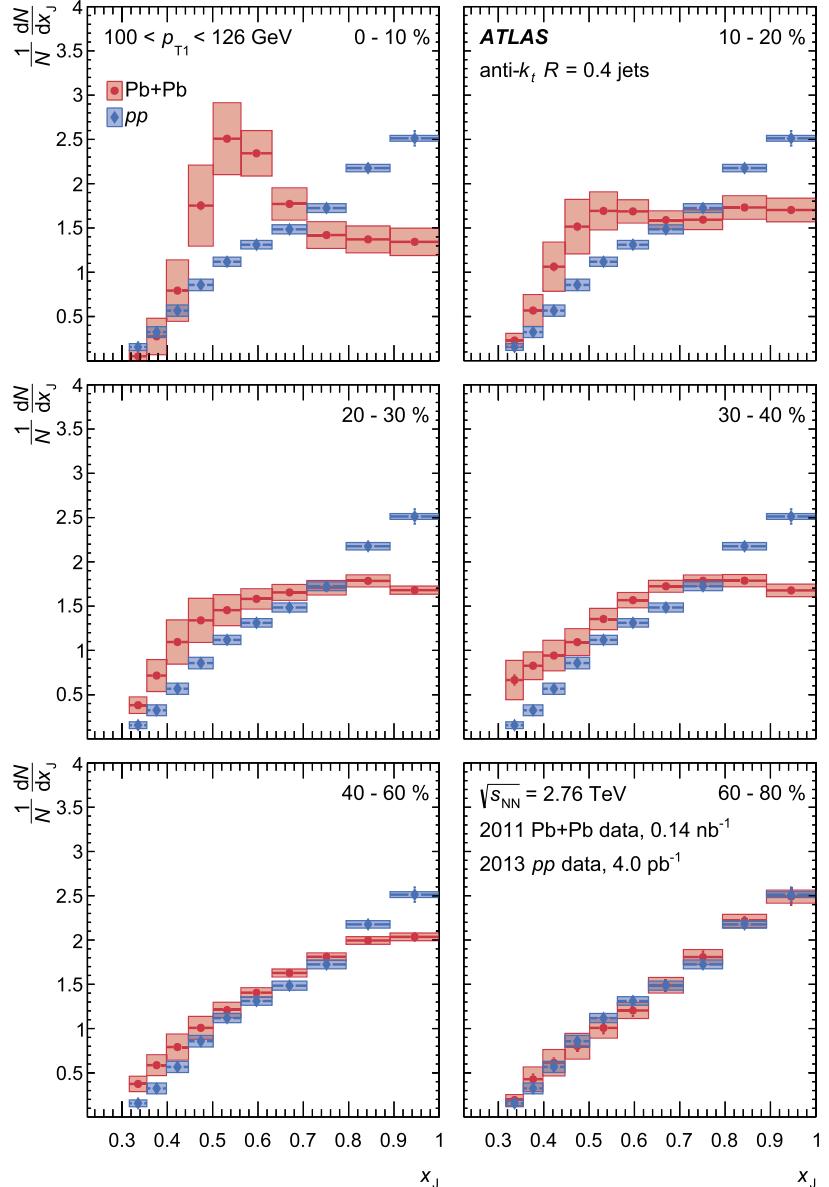


Fig. 8. The $(1/N)dN/dx_J$ distributions for jet pairs with $100 < p_{T1} < 126$ GeV for different collision centralities for $R = 0.4$ jets. The Pb+Pb data are shown in red circles, while the pp distribution is shown for comparison in blue diamonds, and is the same in all panels. Statistical uncertainties are indicated by the error bars while systematic uncertainties are shown with shaded boxes. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

Fig. 9 shows the $(1/N)dN/dx_J$ distributions for 0–10% centrality Pb+Pb collisions and pp collisions for different selections on p_{T1} . In pp collisions, the x_J distribution becomes increasingly narrow with increasing p_{T1} , indicating that higher- p_T dijets tend to be better balanced in momentum (fractionally). At higher p_{T1} , the x_J distribution begins to fall more steeply from $x_J \sim 1$, but appears to flatten at intermediate values of x_J . The modifications observed in the Pb+Pb data lessen with increasing p_{T1} and for jet pairs with $p_{T1} > 200$ GeV the maximum at $x_J \sim 1$ is restored.

The distributions for $R = 0.3$ jets are also shown for the 0–10% centrality interval and for pp collisions for different p_{T1} ranges in **Fig. 10**. The p_T of an $R = 0.3$ jet is generally lower than that of an $R = 0.4$ jet originating from the same hard scattering, and thus features observed in the $(1/N)dN/dx_J$ distributions for $R = 0.4$ jets are expected to appear at lower values of p_{T1} for $R = 0.3$ jets. To facilitate a comparison between results obtained with the two R values, the $R = 0.3$ jet results include an additional p_{T1} interval, $79 < p_{T1} < 100$ GeV. The differences between the Pb+Pb

and pp $(1/N)dN/dx_J$ distributions are qualitatively similar to those observed for $R = 0.4$ jets. **Fig. 11** shows the $(1/N)dN/dx_J$ distributions for $79 < p_{T1} < 100$ GeV for different collision centralities but for jets reconstructed with $R = 0.3$. This indicates that the trends present in p_{T1} and centrality are robust with respect to the UE and that UE effects are properly accounted for by the combinatoric subtraction and unfolding procedures applied in the data analysis. The distributions are flatter for $R = 0.3$ jets in all p_T and centrality bins, including in pp collisions. This is consistent with the expectation that the (p_{T1}, p_{T2}) correlation is weaker for smaller- R jets due to the effects of parton radiation outside the nominal jet cone.

9. Conclusion

This Letter presents a measurement of dijet x_J distributions in 4.0 pb^{-1} of pp and 0.14 nb^{-1} of Pb+Pb collisions at $\sqrt{s_{\text{NN}}} = 2.76 \text{ TeV}$. The measurement is performed differentially in leading-jet transverse momentum, p_{T1} , and in collision centrality using

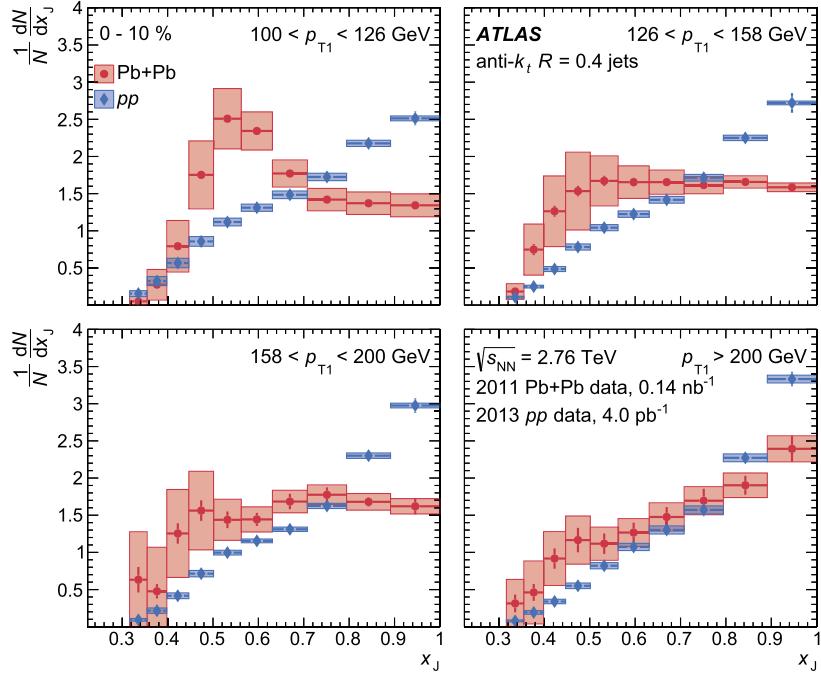


Fig. 9. The $(1/N)dN/dx_J$ distributions for $R = 0.4$ jets with different selections on p_{T1} , shown for the 0–10% centrality bin (red circles) and for pp (blue diamonds). Statistical uncertainties are indicated by the error bars while systematic uncertainties are shown with shaded boxes. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

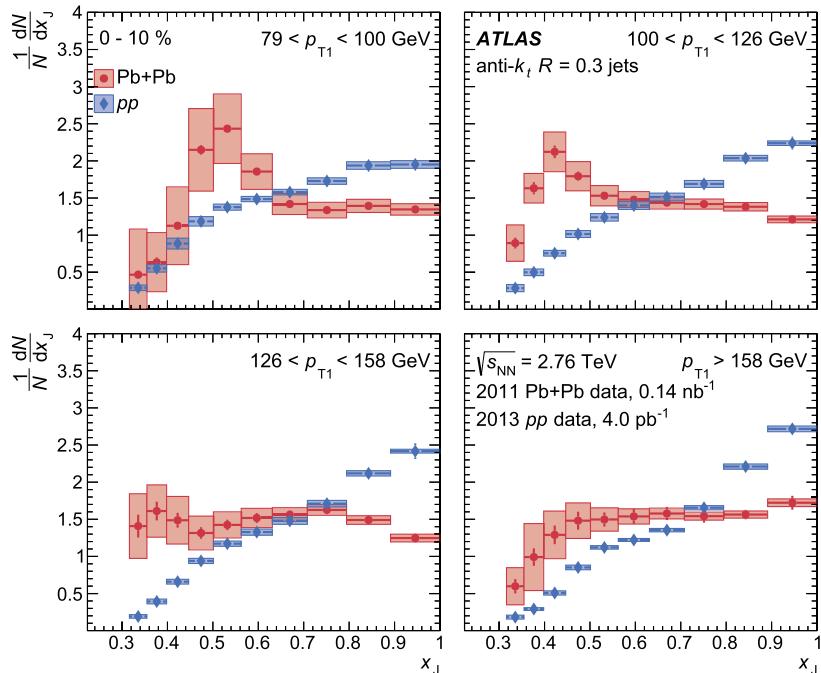


Fig. 10. The $(1/N)dN/dx_J$ distributions for $R = 0.3$ jets with different selections on p_{T1} , shown for the 0–10% centrality bin (red circles) and for pp (blue diamonds). Statistical uncertainties are indicated by the error bars while systematic uncertainties are shown with shaded boxes. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

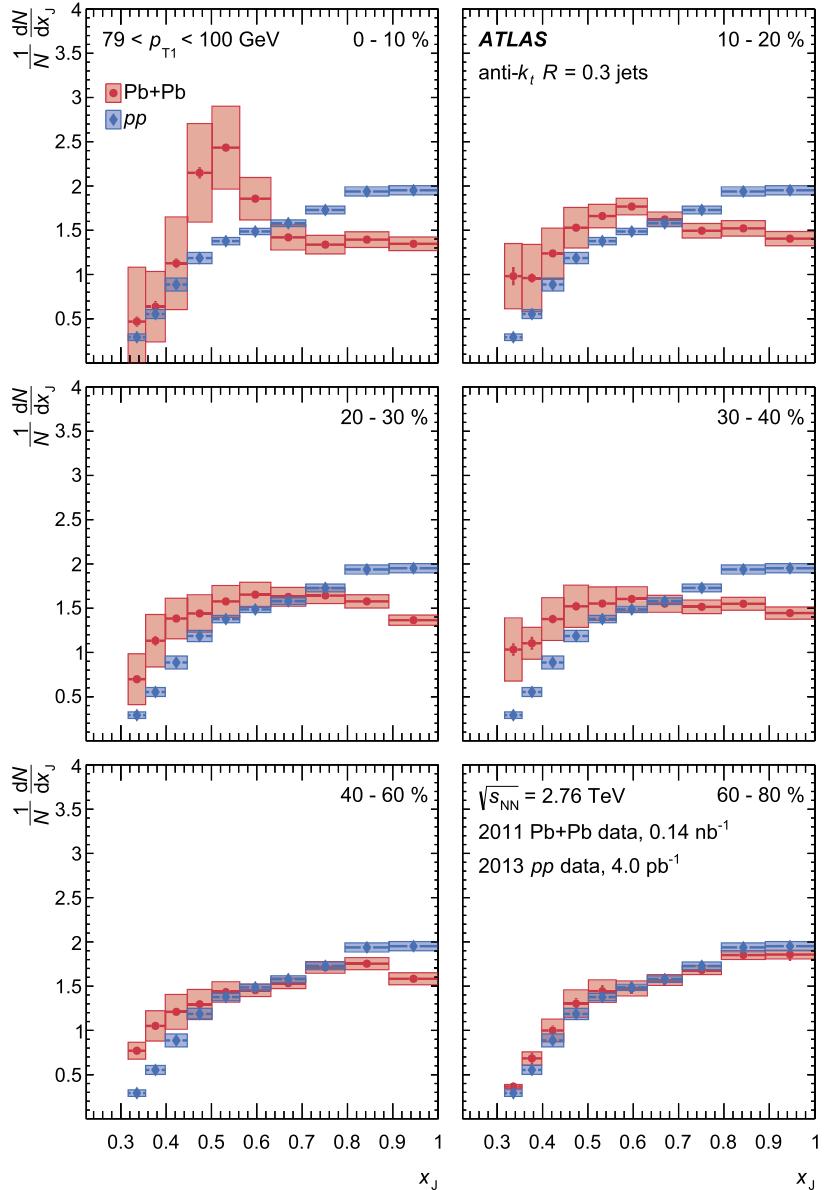


Fig. 11. The $(1/N)dN/dx_J$ distributions for jet pairs with $79 < p_{T_1} < 100$ GeV for different collision centralities for $R = 0.3$ jets. The Pb+Pb data are shown in red circles, while the pp distribution is shown for comparison in blue diamonds, and is the same in all panels. Statistical uncertainties are indicated by the error bars while systematic uncertainties are shown with shaded boxes. (For interpretation of the references to colour in this figure, the reader is referred to the web version of this article.)

data from the ATLAS detector at the LHC. The measured distributions are unfolded to account for the effects of experimental resolution and inefficiencies on the two-dimensional (p_{T_1}, p_{T_2}) distributions and then projected into bins of fixed ratio $x_J = p_{T_2}/p_{T_1}$. The distributions show a larger contribution of asymmetric dijets in Pb+Pb data compared to that in pp data, a feature that becomes more pronounced in more central collisions and is consistent with expectations of medium-induced energy loss due to jet quenching. In the 0–10% centrality bin for $100 < p_{T_1} < 126$ GeV, the x_J distribution develops a significant peak at $x_J \sim 0.5$ indicating that the most probable configuration for dijets is for them to be highly unbalanced. This is in sharp contrast to the situation in the pp data where the most probable values are near $x_J \sim 1$. The centrality-dependent modifications evolve smoothly from central to peripheral collisions, and the results in the 60–80% centrality bin and the pp data are consistent. At larger values of p_{T_1} the x_J distributions are observed to narrow and the differences between the distributions in central Pb+Pb and pp collisions lessen. This is qualita-

tively consistent with a picture in which the fractional energy loss decreases with increasing jet p_T . The features in the data are compatible with those observed in previous measurements of dijets in Pb+Pb collisions by the ATLAS and CMS collaborations, however, the trends in this measurement are more prominent due to the application of the unfolding procedure. This result constitutes an important benchmark for theoretical models of jet quenching and the dynamics of relativistic heavy-ion collisions.

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References

- [1] PHENIX Collaboration, K. Adcox, et al., Formation of dense partonic matter in relativistic nucleus nucleus collisions at RHIC: experimental evaluation by the PHENIX collaboration, *Nucl. Phys. A* 757 (2005) 184, arXiv:nucl-ex/0410003.
- [2] STAR Collaboration, J. Adams, et al., Experimental and theoretical challenges in the search for the quark gluon plasma: the STAR collaboration's critical assessment of the evidence from RHIC collisions, *Nucl. Phys. A* 757 (2005) 102, arXiv:nucl-ex/0501009.
- [3] ATLAS Collaboration, Observation of a centrality-dependent dijet asymmetry in lead-lead collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector at the LHC, *Phys. Rev. Lett.* 105 (2010) 252303, arXiv:1011.6182 [hep-ex].
- [4] J.D. Bjorken, Energy loss of energetic partons in quark-gluon plasma: possible extinction of high p_T jets in hadron-hadron collisions, FERMILAB-PUB-82-059-THY, 1982.
- [5] ATLAS Collaboration, Measurement of the jet radius and transverse momentum dependence of inclusive jet suppression in lead-lead collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector, *Phys. Lett. B* 719 (2013) 220, arXiv:1208.1967 [hep-ex].
- [6] ATLAS Collaboration, Measurements of the nuclear modification factor for jets in Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector, *Phys. Rev. Lett.* 114 (2015) 072302, arXiv:1411.2357 [hep-ex].
- [7] ATLAS Collaboration, Measurement of the azimuthal angle dependence of inclusive jet yields in Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector, *Phys. Rev. Lett.* 111 (2013) 152301, arXiv:1306.6469 [hep-ex].
- [8] CMS Collaboration, Observation and studies of jet quenching in PbPb collisions at $\sqrt{s_{NN}} = 2.76$ TeV, *Phys. Rev. C* 84 (2011) 024906, arXiv:1102.1957 [hep-ex].
- [9] CMS Collaboration, Jet momentum dependence of jet quenching in PbPb collisions at $\sqrt{s_{NN}} = 2.76$ TeV, *Phys. Lett. B* 712 (2012) 176, arXiv:1202.5022 [hep-ex].
- [10] CMS Collaboration, Studies of jet quenching using isolated-photon+jet correlations in PbPb and pp collisions at $\sqrt{s_{NN}} = 2.76$ TeV, *Phys. Lett. B* 718 (2013) 773, arXiv:1205.0206 [hep-ex].
- [11] ATLAS Collaboration, Measurement of inclusive jet charged-particle fragmentation functions in Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector, *Phys. Lett. B* 739 (2014) 320, arXiv:1406.2979 [hep-ex].
- [12] CMS Collaboration, Measurement of jet fragmentation into charged particles in pp and PbPb collisions at $\sqrt{s_{NN}} = 2.76$ TeV, *J. High Energy Phys.* 10 (2012) 087, arXiv:1205.5872 [hep-ex].
- [13] CMS Collaboration, Measurement of jet fragmentation in PbPb and pp collisions at $\sqrt{s_{NN}} = 2.76$ TeV, *Phys. Rev. C* 90 (2014) 024908, arXiv:1406.0932 [hep-ex].
- [14] ALICE Collaboration, B. Abelev, et al., Measurement of charged jet suppression in Pb-Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV, *J. High Energy Phys.* 03 (2014) 013, arXiv:1311.0633 [nucl-ex].
- [15] M. Cacciari, G.P. Salam, G. Soyez, The anti- k_T jet clustering algorithm, *J. High Energy Phys.* 04 (2008) 063, arXiv:0802.1189 [hep-ph].
- [16] ATLAS Collaboration, The ATLAS experiment at the CERN Large Hadron Collider, *J. Instrum.* 3 (2008) S08003.
- [17] ATLAS Collaboration, Measurement of the pseudorapidity and transverse momentum dependence of the elliptic flow of charged particles in lead-lead collisions at $\sqrt{s_{NN}} = 2.76$ TeV with the ATLAS detector, *Phys. Lett. B* 707 (2012) 330, arXiv:1108.6018 [hep-ex].
- [18] ATLAS Collaboration, The performance of the jet trigger for the ATLAS detector during 2011 data taking, *Eur. Phys. J. C* 76 (2016) 526, arXiv:1606.07759 [hep-ex].
- [19] ATLAS Collaboration, Jet energy measurement with the ATLAS detector in proton-proton collisions at $\sqrt{s} = 7$ TeV, *Eur. Phys. J. C* 73 (2013) 2304, arXiv:1112.6426 [hep-ex].
- [20] M.L. Miller, et al., Glauber modeling in high energy nuclear collisions, *Annu. Rev. Nucl. Part. Sci.* 57 (2007) 205, arXiv:nucl-ex/0701025.
- [21] B. Alver, et al., The PHOBOS Glauber Monte Carlo, arXiv:0805.4411 [nucl-ex], 2008.
- [22] S. Agostinelli, et al., GEANT4: a simulation toolkit, *Nucl. Instrum. Methods Phys. Res., Sect. A, Accel. Spectrom. Detect. Assoc. Equip.* 506 (2003) 250.
- [23] ATLAS Collaboration, The ATLAS simulation infrastructure, *Eur. Phys. J. C* 70 (2010) 823, arXiv:1005.4568 [hep-ex].
- [24] T. Sjöstrand, S. Mrenna, P.Z. Skands, PYTHIA 6.4 physics and manual, *J. High Energy Phys.* 05 (2006) 026, arXiv:hep-ph/0603175.
- [25] ATLAS Collaboration, ATLAS tunes of PYTHIA 6 and Pythia 8 for MC11, ATL-PHYS-PUB-2011-009, <http://cdsweb.cern.ch/record/1363300>, 2011.
- [26] J. Pumplin, et al., New generation of parton distributions with uncertainties from global QCD analysis, *J. High Energy Phys.* 07 (2002) 012, arXiv:hep-ph/0201195.
- [27] T. Sjöstrand, S. Mrenna, P.Z. Skands, A brief introduction to PYTHIA 8.1, *Comput. Phys. Commun.* 178 (2008) 852, arXiv:0710.3820 [hep-ph].
- [28] ATLAS Collaboration, Summary of ATLAS Pythia 8 tunes, ATL-PHYS-PUB-2012-003, <https://cds.cern.ch/record/1474107>, 2012.
- [29] H.-L. Lai, et al., New parton distributions for collider physics, *Phys. Rev. D* 82 (2010) 074024, arXiv:1007.2241 [hep-ph].
- [30] I. Dokshitzer, A. Snigirev, A model of jet quenching in ultrarelativistic heavy ion collisions and high- $p(T)$ hadron spectra at RHIC, *Eur. Phys. J. C* 45 (2006) 211, arXiv:hep-ph/0506189.
- [31] M. Cacciari, G.P. Salam, G. Soyez, FastJet user manual, *Eur. Phys. J. C* 72 (2012) 1896, arXiv:1111.6097 [hep-ph].
- [32] ATLAS Collaboration, Jet energy measurement and its systematic uncertainty in proton-proton collisions at $\sqrt{s} = 7$ TeV with the ATLAS detector, *Eur. Phys. J. C* 75 (2015) 17, arXiv:1406.0076 [hep-ex].
- [33] ATLAS Collaboration, Measurement of the azimuthal anisotropy for charged particle production in $\sqrt{s_{NN}} = 2.76$ TeV lead-lead collisions with the ATLAS detector, *Phys. Rev. C* 86 (2012) 014907, arXiv:1203.3087 [hep-ex].
- [34] G. D'Agostini, A multidimensional unfolding method based on Bayes' theorem, *Nucl. Instrum. Methods Phys. Res., Sect. A, Accel. Spectrom. Detect. Assoc. Equip.* 362 (1995) 487.
- [35] T. Adye, Unfolding algorithms and tests using RooUnfold, arXiv:1105.1160 [physics.data-an], 2011.
- [36] ATLAS Collaboration, Jet energy resolution in proton-proton collisions at $\sqrt{s} = 7$ TeV recorded in 2010 with the ATLAS detector, *Eur. Phys. J. C* 73 (2013) 2306, arXiv:1210.6210 [hep-ex].
- [37] ATLAS Collaboration, Study of jet shapes in inclusive jet production in pp collisions at $\sqrt{s} = 7$ TeV using the ATLAS detector, *Phys. Rev. D* 83 (2011) 052003, arXiv:1101.0070 [hep-ex].
- [38] M. Bahr, et al., Herwig++ physics and manual, *Eur. Phys. J. C* 58 (2008) 639, arXiv:0803.0883 [hep-ph].
- [39] S. Gieseke, C. Rohr, A. Siodmok, Colour reconnections in Herwig++, *Eur. Phys. J. C* 72 (2012) 2225, arXiv:1206.0041 [hep-ph].
- [40] P. Nason, A new method for combining NLO QCD with shower Monte Carlo algorithms, *J. High Energy Phys.* 11 (2004) 040, arXiv:hep-ph/0409146.
- [41] S. Frixione, P. Nason, C. Oleari, Matching NLO QCD computations with parton shower simulations: the POWHEG method, *J. High Energy Phys.* 11 (2007) 070, arXiv:0709.2092 [hep-ph].
- [42] S. Alioli, et al., A general framework for implementing NLO calculations in shower Monte Carlo programs: the POWHEG BOX, *J. High Energy Phys.* 06 (2010) 043, arXiv:1002.2581 [hep-ph].
- [43] ATLAS Collaboration, ATLAS computing acknowledgements 2016–2017, ATL-GEN-PUB-2016-002, 2016, <https://cdsweb.cern.ch/record/2202407>.

The ATLAS Collaboration

- M. Aaboud ^{137d}, G. Aad ⁸⁸, B. Abbott ¹¹⁵, J. Abdallah ⁸, O. Abdinov ¹², B. Abelsoos ¹¹⁹, S.H. Abidi ¹⁶¹, O.S. AbouZeid ¹³⁹, N.L. Abraham ¹⁵¹, H. Abramowicz ¹⁵⁵, H. Abreu ¹⁵⁴, R. Abreu ¹¹⁸, Y. Abulaiti ^{148a,148b}, B.S. Acharya ^{167a,167b,a}, S. Adachi ¹⁵⁷, L. Adamczyk ^{41a}, D.L. Adams ²⁷, J. Adelman ¹¹⁰, M. Adersberger ¹⁰², T. Adye ¹³³, A.A. Affolder ¹³⁹, T. Agatonovic-Jovin ¹⁴, C. Agheorghiesei ^{28b}, J.A. Aguilar-Saavedra ^{128a,128f}, S.P. Ahlen ²⁴, F. Ahmadov ^{68,b}, G. Aielli ^{135a,135b}, S. Akatsuka ⁷¹, H. Akerstedt ^{148a,148b}, T.P.A. Åkesson ⁸⁴, A.V. Akimov ⁹⁸, G.L. Alberghi ^{22a,22b}, J. Albert ¹⁷², M.J. Alconada Verzini ⁷⁴, M. Aleksić ³², I.N. Aleksandrov ⁶⁸, C. Alexa ^{28b}, G. Alexander ¹⁵⁵, T. Alexopoulos ¹⁰, M. Alhroob ¹¹⁵, B. Ali ¹³⁰, M. Aliev ^{76a,76b}, G. Alimonti ^{94a}, J. Alison ³³, S.P. Alkire ³⁸, B.M.M. Allbrooke ¹⁵¹, B.W. Allen ¹¹⁸, P.P. Allport ¹⁹, A. Aloisio ^{106a,106b}, A. Alonso ³⁹, F. Alonso ⁷⁴, C. Alpigiani ¹⁴⁰, A.A. Alshehri ⁵⁶, M. 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Asbah ⁴⁵, A. Ashkenazi ¹⁵⁵, L. Asquith ¹⁵¹, K. Assamagan ²⁷, R. Astalos ^{146a}, M. Atkinson ¹⁶⁹, N.B. Atlay ¹⁴³, K. Augsten ¹³⁰, G. Avolio ³², B. Axen ¹⁶, M.K. Ayoub ¹¹⁹, G. Azuelos ^{97,d}, A.E. Baas ^{60a}, M.J. Baca ¹⁹, H. Bachacou ¹³⁸, K. Bachas ^{76a,76b}, M. Backes ¹²², M. Backhaus ³², P. Bagiacchi ^{134a,134b}, P. Bagnaia ^{134a,134b}, J.T. Baines ¹³³, M. Bajic ³⁹, O.K. Baker ¹⁷⁹, E.M. Baldin ^{111,c}, P. Balek ¹⁷⁵, T. Balestri ¹⁵⁰, F. Balli ¹³⁸, W.K. Balunas ¹²⁴, E. Banas ⁴², Sw. Banerjee ^{176,e}, A.A.E. Bannoura ¹⁷⁸, L. Barak ³², E.L. Barberio ⁹¹, D. Barberis ^{53a,53b}, M. Barbero ⁸⁸, T. Barillari ¹⁰³, M-S Barisits ³², T. Barklow ¹⁴⁵, N. Barlow ³⁰, S.L. Barnes ^{36c}, B.M. Barnett ¹³³, R.M. Barnett ¹⁶, Z. Barnovska-Blenessy ^{36a}, A. Baroncelli ^{136a}, G. Barone ²⁵, A.J. Barr ¹²², L. Barranco Navarro ¹⁷⁰, F. Barreiro ⁸⁵, J. Barreiro Guimarães da Costa ^{35a}, R. Bartoldus ¹⁴⁵, A.E. Barton ⁷⁵, P. Bartos ^{146a}, A. Basalaev ¹²⁵, A. Bassalat ^{119,f}, R.L. Bates ⁵⁶, S.J. Batista ¹⁶¹, J.R. Batley ³⁰, M. Battaglia ¹³⁹, M. Bauce ^{134a,134b}, F. Bauer ¹³⁸, H.S. Bawa ^{145,g}, J.B. Beacham ¹¹³, M.D. Beattie ⁷⁵, T. Beau ⁸³, P.H. Beauchemin ¹⁶⁵, P. Bechtle ²³, H.P. Beck ^{18,h}, K. Becker ¹²², M. Becker ⁸⁶, M. Beckingham ¹⁷³, C. Becot ¹¹², A.J. Beddall ^{20d}, A. Beddall ^{20b}, V.A. Bednyakov ⁶⁸, M. Bedognetti ¹⁰⁹, C.P. Bee ¹⁵⁰, T.A. Beermann ³², M. Begalli ^{26a}, M. Begel ²⁷, J.K. Behr ⁴⁵, A.S. Bell ⁸¹, G. Bella ¹⁵⁵, L. Bellagamba ^{22a}, A. Bellerive ³¹, M. Bellomo ⁸⁹, K. Belotskiy ¹⁰⁰, O. Beltramello ³², N.L. Belyaev ¹⁰⁰, O. Benary ^{155,*}, D. Benchekroun ^{137a}, M. Bender ¹⁰², K. Bendtz ^{148a,148b}, N. Benekos ¹⁰, Y. Benhammou ¹⁵⁵, E. Benhar Noccioli ¹⁷⁹, J. Benitez ⁶⁶, D.P. Benjamin ⁴⁸, M. Benoit ⁵², J.R. Bensinger ²⁵, S. Bentvelsen ¹⁰⁹, L. Beresford ¹²², M. Beretta ⁵⁰, D. Berge ¹⁰⁹, E. Bergeaas Kuutmann ¹⁶⁸, N. Berger ⁵, J. Beringer ¹⁶, S. Berlendis ⁵⁸, N.R. Bernard ⁸⁹, G. Bernardi ⁸³, C. Bernius ¹¹², F.U. Bernlochner ²³, T. Berry ⁸⁰, P. Berta ¹³¹, C. Bertella ⁸⁶, G. Bertoli ^{148a,148b}, F. Bertolucci ^{126a,126b}, I.A. Bertram ⁷⁵, C. Bertsche ⁴⁵, D. Bertsche ¹¹⁵, G.J. Besjes ³⁹, O. Bessidskaia Bylund ^{148a,148b}, M. Bessner ⁴⁵, N. Besson ¹³⁸, C. Betancourt ⁵¹, A. Bethani ⁸⁷, S. Bethke ¹⁰³, A.J. Bevan ⁷⁹, R.M. Bianchi ¹²⁷, M. Bianco ³², O. Biebel ¹⁰², D. Biedermann ¹⁷, R. Bielski ⁸⁷, N.V. Biesuz ^{126a,126b}, M. Biglietti ^{136a}, J. Bilbao De Mendizabal ⁵², T.R.V. Billoud ⁹⁷, H. Bilokon ⁵⁰, M. Bindi ⁵⁷, A. Bingul ^{20b}, C. Bini ^{134a,134b}, S. Biondi ^{22a,22b}, T. Bisanz ⁵⁷, C. Bittrich ⁴⁷, D.M. Bjergaard ⁴⁸, C.W. Black ¹⁵², J.E. Black ¹⁴⁵, K.M. Black ²⁴, D. Blackburn ¹⁴⁰, R.E. Blair ⁶, T. Blazek ^{146a}, I. Bloch ⁴⁵, C. Blocker ²⁵, A. Blue ⁵⁶, W. Blum ^{86,*}, U. Blumenschein ⁷⁹, S. Blunier ^{34a}, G.J. Bobbink ¹⁰⁹, V.S. Bobrovnikov ^{111,c}, S.S. Bocchetta ⁸⁴, A. Bocci ⁴⁸, C. Bock ¹⁰², M. Boehler ⁵¹, D. Boerner ¹⁷⁸, D. Bogavac ¹⁰², A.G. Bogdanchikov ¹¹¹, C. Bohm ^{148a}, V. Boisvert ⁸⁰, P. Bokan ^{168,i}, T. Bold ^{41a}, A.S. Boldyrev ¹⁰¹, M. Bomben ⁸³, M. Bona ⁷⁹, M. Boonekamp ¹³⁸, A. Borisov ¹³², G. Borissov ⁷⁵, J. Bortfeldt ³², D. Bortoletto ¹²², V. Bortolotto ^{62a,62b,62c}, D. Boscherini ^{22a}, M. Bosman ¹³, J.D. Bossio Sola ²⁹, J. Boudreau ¹²⁷, J. Bouffard ², E.V. Bouhova-Thacker ⁷⁵, D. Boumediene ³⁷, C. Bourdarios ¹¹⁹, S.K. Boutle ⁵⁶, A. Boveia ¹¹³, J. Boyd ³², I.R. Boyko ⁶⁸, J. Bracinik ¹⁹, A. Brandt ⁸, G. Brandt ⁵⁷, O. Brandt ^{60a}, U. Bratzler ¹⁵⁸, B. Brau ⁸⁹, J.E. Brau ¹¹⁸, W.D. Breaden Madden ⁵⁶, K. Brendlinger ⁴⁵, A.J. Brennan ⁹¹, L. Brenner ¹⁰⁹, R. Brenner ¹⁶⁸, S. Bressler ¹⁷⁵, D.L. Briglin ¹⁹,

- T.M. Bristow 49, D. Britton 56, D. Blitzger 45, F.M. Brochu 30, I. Brock 23, R. Brock 93, G. Brooijmans 38,
 T. Brooks 80, W.K. Brooks 34b, J. Brosamer 16, E. Brost 110, J.H. Broughton 19, P.A. Bruckman de Renstrom 42,
 D. Bruncko 146b, A. Bruni 22a, G. Bruni 22a, L.S. Bruni 109, B.H. Brunt 30, M. Bruschi 22a, N. Bruscino 23,
 P. Bryant 33, L. Bryngemark 84, T. Buanes 15, Q. Buat 144, P. Buchholz 143, A.G. Buckley 56, I.A. Budagov 68,
 F. Buehrer 51, M.K. Bugge 121, O. Bulekov 100, D. Bullock 8, H. Burckhart 32, S. Burdin 77, C.D. Burgard 51,
 A.M. Burger 5, B. Burghgrave 110, K. Burka 42, S. Burke 133, I. Burmeister 46, J.T.P. Burr 122, E. Busato 37,
 D. Büscher 51, V. Büscher 86, P. Bussey 56, J.M. Butler 24, C.M. Buttar 56, J.M. Butterworth 81, P. Butti 32,
 W. Buttlinger 27, A. Buzatu 35c, A.R. Buzykaev 111,c, S. Cabrera Urbán 170, D. Caforio 130, V.M. Cairo 40a,40b,
 O. Cakir 4a, N. Calace 52, P. Calafiura 16, A. Calandri 88, G. Calderini 83, P. Calfayan 64, G. Callea 40a,40b,
 L.P. Caloba 26a, S. Calvente Lopez 85, D. Calvet 37, S. Calvet 37, T.P. Calvet 88, R. Camacho Toro 33,
 S. Camarda 32, P. Camarri 135a,135b, D. Cameron 121, R. Caminal Armadans 169, C. Camincher 58,
 S. Campana 32, M. Campanelli 81, A. Camplani 94a,94b, A. Campoverde 143, V. Canale 106a,106b,
 M. Cano Bret 36c, J. Cantero 116, T. Cao 155, M.D.M. Capeans Garrido 32, I. Caprini 28b, M. Caprini 28b,
 M. Capua 40a,40b, R.M. Carbone 38, R. Cardarelli 135a, F. Cardillo 51, I. Carli 131, T. Carli 32, G. Carlino 106a,
 B.T. Carlson 127, L. Carminati 94a,94b, R.M.D. Carney 148a,148b, S. Caron 108, E. Carquin 34b,
 G.D. Carrillo-Montoya 32, J. Carvalho 128a,128c, D. Casadei 19, M.P. Casado 13,j, M. Casolino 13,
 D.W. Casper 166, R. Castelijn 109, A. Castelli 109, V. Castillo Gimenez 170, N.F. Castro 128a,k, A. Catinaccio 32,
 J.R. Catmore 121, A. Cattai 32, J. Caudron 23, V. Cavalieri 169, E. Cavallaro 13, D. Cavalli 94a,
 M. Cavalli-Sforza 13, V. Cavasinni 126a,126b, E. Celebi 20a, F. Ceradini 136a,136b, L. Cerdá Alberich 170,
 A.S. Cerqueira 26b, A. Cerri 151, L. Cerrito 135a,135b, F. Cerutti 16, A. Cervelli 18, S.A. Cetin 20c, A. Chafaq 137a,
 D. Chakraborty 110, S.K. Chan 59, W.S. Chan 109, Y.L. Chan 62a, P. Chang 169, J.D. Chapman 30,
 D.G. Charlton 19, A. Chatterjee 52, C.C. Chau 161, C.A. Chavez Barajas 151, S. Che 113, S. Cheatham 167a,167c,
 A. Chegwidden 93, S. Chekanov 6, S.V. Chekulaev 163a, G.A. Chelkov 68,l, M.A. Chelstowska 32, C. Chen 67,
 H. Chen 27, S. Chen 35b, S. Chen 157, X. Chen 35c,m, Y. Chen 70, H.C. Cheng 92, H.J. Cheng 35a, Y. Cheng 33,
 A. Cheplakov 68, E. Cheremushkina 132, R. Cherkaoui El Moursli 137e, V. Chernyatin 27,* , E. Cheu 7,
 L. Chevalier 138, V. Chiarella 50, G. Chiarella 126a,126b, G. Chiodini 76a, A.S. Chisholm 32, A. Chitan 28b,
 Y.H. Chiu 172, M.V. Chizhov 68, K. Choi 64, A.R. Chomont 37, S. Chouridou 9, B.K.B. Chow 102,
 V. Christodoulou 81, D. Chromek-Burckhart 32, M.C. Chu 62a, J. Chudoba 129, A.J. Chuinard 90,
 J.J. Chwastowski 42, L. Chytka 117, A.K. Ciftci 4a, D. Cinca 46, V. Cindro 78, I.A. Cioara 23, C. Ciocca 22a,22b,
 A. Ciocio 16, F. Cirotto 106a,106b, Z.H. Citron 175, M. Citterio 94a, M. Ciubancan 28b, A. Clark 52, B.L. Clark 59,
 M.R. Clark 38, P.J. Clark 49, R.N. Clarke 16, C. Clement 148a,148b, Y. Coadou 88, M. Cobal 167a,167c,
 A. Coccaro 52, J. Cochran 67, L. Colasurdo 108, B. Cole 38, A.P. Colijn 109, J. Collot 58, T. Colombo 166,
 P. Conde Muiño 128a,128b, E. Coniavitis 51, S.H. Connell 147b, I.A. Connelly 87, V. Consorti 51,
 S. Constantinescu 28b, G. Conti 32, F. Conventi 106a,n, M. Cooke 16, B.D. Cooper 81, A.M. Cooper-Sarkar 122,
 F. Cormier 171, K.J.R. Cormier 161, T. Cornelissen 178, M. Corradi 134a,134b, F. Corriveau 90,o,
 A. Cortes-Gonzalez 32, G. Cortiana 103, G. Costa 94a, M.J. Costa 170, D. Costanzo 141, G. Cottin 30,
 G. Cowan 80, B.E. Cox 87, K. Cranmer 112, S.J. Crawley 56, R.A. Creager 124, G. Cree 31, S. Crépé-Renaudin 58,
 F. Crescioli 83, W.A. Cribbs 148a,148b, M. Crispin Ortuzar 122, M. Cristinziani 23, V. Croft 108,
 G. Crosetti 40a,40b, A. Cueto 85, T. Cuhadar Donszelmann 141, J. Cummings 179, M. Curatolo 50, J. Cúth 86,
 H. Czirr 143, P. Czodrowski 32, G. D'amen 22a,22b, S. D'Auria 56, M. D'Onofrio 77,
 M.J. Da Cunha Sargedas De Sousa 128a,128b, C. Da Via 87, W. Dabrowski 41a, T. Dado 146a, T. Dai 92,
 O. Dale 15, F. Dallaire 97, C. Dallapiccola 89, M. Dam 39, J.R. Dandoy 124, N.P. Dang 51, A.C. Daniells 19,
 N.S. Dann 87, M. Danninger 171, M. Dano Hoffmann 138, V. Dao 150, G. Darbo 53a, S. Darmora 8,
 J. Dassoulas 3, A. Dattagupta 118, T. Daubney 45, W. Davey 23, C. David 45, T. Davidek 131, M. Davies 155,
 P. Davison 81, E. Dawe 91, I. Dawson 141, K. De 8, R. de Asmundis 106a, A. De Benedetti 115,
 S. De Castro 22a,22b, S. De Cecco 83, N. De Groot 108, P. de Jong 109, H. De la Torre 93, F. De Lorenzi 67,
 A. De Maria 57, D. De Pedis 134a, A. De Salvo 134a, U. De Sanctis 151, A. De Santo 151,
 K. De Vasconcelos Corga 88, J.B. De Vivie De Regie 119, W.J. Dearnaley 75, R. Debbe 27, C. Debenedetti 139,
 D.V. Dedovich 68, N. Dehghanian 3, I. Deigaard 109, M. Del Gaudio 40a,40b, J. Del Peso 85,
 T. Del Prete 126a,126b, D. Delgove 119, F. Deliot 138, C.M. Delitzsch 52, A. Dell'Acqua 32, L. Dell'Asta 24,
 M. Dell'Orso 126a,126b, M. Della Pietra 106a,106b, D. della Volpe 52, M. Delmastro 5, P.A. Delsart 58,
 D.A. DeMarco 161, S. Demers 179, M. Demichev 68, A. Demilly 83, S.P. Denisov 132, D. Denysiuk 138,

- D. Derendarz ⁴², J.E. Derkaoui ^{137d}, F. Derue ⁸³, P. Dervan ⁷⁷, K. Desch ²³, C. Deterre ⁴⁵, K. Dette ⁴⁶,
 P.O. Deviveiros ³², A. Dewhurst ¹³³, S. Dhaliwal ²⁵, A. Di Ciaccio ^{135a,135b}, L. Di Ciaccio ⁵,
 W.K. Di Clemente ¹²⁴, C. Di Donato ^{106a,106b}, A. Di Girolamo ³², B. Di Girolamo ³², B. Di Micco ^{136a,136b},
 R. Di Nardo ³², K.F. Di Petrillo ⁵⁹, A. Di Simone ⁵¹, R. Di Sipio ¹⁶¹, D. Di Valentino ³¹, C. Diaconu ⁸⁸,
 M. Diamond ¹⁶¹, F.A. Dias ⁴⁹, M.A. Diaz ^{34a}, E.B. Diehl ⁹², J. Dietrich ¹⁷, S. Díez Cornell ⁴⁵,
 A. Dimitrijevska ¹⁴, J. Dingfelder ²³, P. Dita ^{28b}, S. Dita ^{28b}, F. Dittus ³², F. Djama ⁸⁸, T. Djebava ^{54b},
 J.I. Djuvstrand ^{60a}, M.A.B. do Vale ^{26c}, D. Dobos ³², M. Dobre ^{28b}, C. Doglioni ⁸⁴, J. Dolejsi ¹³¹, Z. Dolezal ¹³¹,
 M. Donadelli ^{26d}, S. Donati ^{126a,126b}, P. Dondero ^{123a,123b}, J. Donini ³⁷, J. Dopke ¹³³, A. Doria ^{106a},
 M.T. Dova ⁷⁴, A.T. Doyle ⁵⁶, E. Drechsler ⁵⁷, M. Dris ¹⁰, Y. Du ^{36b}, J. Duarte-Campderros ¹⁵⁵, E. Duchovni ¹⁷⁵,
 G. Duckeck ¹⁰², O.A. Ducu ^{97,p}, D. Duda ¹⁰⁹, A. Dudarev ³², A. Chr. Dudder ⁸⁶, E.M. Duffield ¹⁶, L. Duflot ¹¹⁹,
 M. Dührssen ³², M. Dumancic ¹⁷⁵, A.E. Dumitriu ^{28b}, A.K. Duncan ⁵⁶, M. Dunford ^{60a}, H. Duran Yildiz ^{4a},
 M. Düren ⁵⁵, A. Durglishvili ^{54b}, D. Duschinger ⁴⁷, B. Dutta ⁴⁵, M. Dyndal ⁴⁵, C. Eckardt ⁴⁵, K.M. Ecker ¹⁰³,
 R.C. Edgar ⁹², T. Eifert ³², G. Eigen ¹⁵, K. Einsweiler ¹⁶, T. Ekelof ¹⁶⁸, M. El Kacimi ^{137c}, V. Ellajosyula ⁸⁸,
 M. Ellert ¹⁶⁸, S. Elles ⁵, F. Ellinghaus ¹⁷⁸, A.A. Elliot ¹⁷², N. Ellis ³², J. Elmsheuser ²⁷, M. Elsing ³²,
 D. Emeliyanov ¹³³, Y. Enari ¹⁵⁷, O.C. Endner ⁸⁶, J.S. Ennis ¹⁷³, J. Erdmann ⁴⁶, A. Ereditato ¹⁸, G. Ernis ¹⁷⁸,
 M. Ernst ²⁷, S. Errede ¹⁶⁹, E. Ertel ⁸⁶, M. Escalier ¹¹⁹, H. Esch ⁴⁶, C. Escobar ¹²⁷, B. Esposito ⁵⁰,
 A.I. Etienne ¹³⁸, E. Etzion ¹⁵⁵, H. Evans ⁶⁴, A. Ezhilov ¹²⁵, F. Fabbri ^{22a,22b}, L. Fabbri ^{22a,22b}, G. Facini ³³,
 R.M. Fakhrutdinov ¹³², S. Falciano ^{134a}, R.J. Falla ⁸¹, J. Faltova ³², Y. Fang ^{35a}, M. Fanti ^{94a,94b}, A. Farbin ⁸,
 A. Farilla ^{136a}, C. Farina ¹²⁷, E.M. Farina ^{123a,123b}, T. Farooque ⁹³, S. Farrell ¹⁶, S.M. Farrington ¹⁷³,
 P. Farthouat ³², F. Fassi ^{137e}, P. Fassnacht ³², D. Fassouliotis ⁹, M. Facci Giannelli ⁸⁰, A. Favareto ^{53a,53b},
 W.J. Fawcett ¹²², L. Fayard ¹¹⁹, O.L. Fedin ^{125,q}, W. Fedorko ¹⁷¹, S. Feigl ¹²¹, L. Feligioni ⁸⁸, C. Feng ^{36b},
 E.J. Feng ³², H. Feng ⁹², A.B. Fenyuk ¹³², L. Feremenga ⁸, P. Fernandez Martinez ¹⁷⁰, S. Fernandez Perez ¹³,
 J. Ferrando ⁴⁵, A. Ferrari ¹⁶⁸, P. Ferrari ¹⁰⁹, R. Ferrari ^{123a}, D.E. Ferreira de Lima ^{60b}, A. Ferrer ¹⁷⁰,
 D. Ferrere ⁵², C. Ferretti ⁹², F. Fiedler ⁸⁶, A. Filipčič ⁷⁸, M. Filipuzzi ⁴⁵, F. Filthaut ¹⁰⁸, M. Fincke-Keeler ¹⁷²,
 K.D. Finelli ¹⁵², M.C.N. Fiolhais ^{128a,128c,r}, L. Fiorini ¹⁷⁰, A. Fischer ², C. Fischer ¹³, J. Fischer ¹⁷⁸,
 W.C. Fisher ⁹³, N. Flaschel ⁴⁵, I. Fleck ¹⁴³, P. Fleischmann ⁹², R.R.M. Fletcher ¹²⁴, T. Flick ¹⁷⁸, B.M. Flierl ¹⁰²,
 L.R. Flores Castillo ^{62a}, M.J. Flowerdew ¹⁰³, G.T. Forcolin ⁸⁷, A. Formica ¹³⁸, A. Forti ⁸⁷, A.G. Foster ¹⁹,
 D. Fournier ¹¹⁹, H. Fox ⁷⁵, S. Fracchia ¹³, P. Francavilla ⁸³, M. Franchini ^{22a,22b}, D. Francis ³², L. Franconi ¹²¹,
 M. Franklin ⁵⁹, M. Frate ¹⁶⁶, M. Fraternali ^{123a,123b}, D. Freeborn ⁸¹, S.M. Fressard-Batraneanu ³²,
 B. Freund ⁹⁷, D. Froidevaux ³², J.A. Frost ¹²², C. Fukunaga ¹⁵⁸, E. Fullana Torregrosa ⁸⁶, T. Fusayasu ¹⁰⁴,
 J. Fuster ¹⁷⁰, C. Gabaldon ⁵⁸, O. Gabizon ¹⁵⁴, A. Gabrielli ^{22a,22b}, A. Gabrielli ¹⁶, G.P. Gach ^{41a}, S. Gadatsch ³²,
 S. Gadomski ⁸⁰, G. Gagliardi ^{53a,53b}, L.G. Gagnon ⁹⁷, P. Gagnon ⁶⁴, C. Galea ¹⁰⁸, B. Galhardo ^{128a,128c},
 E.J. Gallas ¹²², B.J. Gallop ¹³³, P. Gallus ¹³⁰, G. Galster ³⁹, K.K. Gan ¹¹³, S. Ganguly ³⁷, J. Gao ^{36a}, Y. Gao ⁷⁷,
 Y.S. Gao ^{145,g}, F.M. Garay Walls ⁴⁹, C. García ¹⁷⁰, J.E. García Navarro ¹⁷⁰, M. Garcia-Sciveres ¹⁶,
 R.W. Gardner ³³, N. Garelli ¹⁴⁵, V. Garonne ¹²¹, A. Gascon Bravo ⁴⁵, K. Gasnikova ⁴⁵, C. Gatti ⁵⁰,
 A. Gaudiello ^{53a,53b}, G. Gaudio ^{123a}, I.L. Gavrilenko ⁹⁸, C. Gay ¹⁷¹, G. Gaycken ²³, E.N. Gazis ¹⁰,
 C.N.P. Gee ¹³³, M. Geisen ⁸⁶, M.P. Geisler ^{60a}, K. Gellerstedt ^{148a,148b}, C. Gemme ^{53a}, M.H. Genest ⁵⁸,
 C. Geng ^{36a,s}, S. Gentile ^{134a,134b}, C. Gentsos ¹⁵⁶, S. George ⁸⁰, D. Gerbaudo ¹³, A. Gershon ¹⁵⁵,
 S. Ghasemi ¹⁴³, M. Ghneimat ²³, B. Giacobbe ^{22a}, S. Giagu ^{134a,134b}, P. Giannetti ^{126a,126b}, S.M. Gibson ⁸⁰,
 M. Gignac ¹⁷¹, M. Gilchriese ¹⁶, D. Gillberg ³¹, G. Gilles ¹⁷⁸, D.M. Gingrich ^{3,d}, N. Giokaris ^{9,*},
 M.P. Giordani ^{167a,167c}, F.M. Giorgi ^{22a}, P.F. Giraud ¹³⁸, P. Giromini ⁵⁹, D. Giugni ^{94a}, F. Giulia ¹²²,
 C. Giuliani ¹⁰³, M. Giulini ^{60b}, B.K. Gjelsten ¹²¹, S. Gkaitatzis ¹⁵⁶, I. Gkialas ^{9,t}, E.L. Gkougkousis ¹³⁹,
 L.K. Gladilin ¹⁰¹, C. Glasman ⁸⁵, J. Glatzer ¹³, P.C.F. Glaysher ⁴⁵, A. Glazov ⁴⁵, M. Goblirsch-Kolb ²⁵,
 J. Godlewski ⁴², S. Goldfarb ⁹¹, T. Golling ⁵², D. Golubkov ¹³², A. Gomes ^{128a,128b,128d}, R. Gonçalo ^{128a},
 R. Goncalves Gama ^{26a}, J. Goncalves Pinto Firmino Da Costa ¹³⁸, G. Gonella ⁵¹, L. Gonella ¹⁹,
 A. Gongadze ⁶⁸, S. González de la Hoz ¹⁷⁰, S. Gonzalez-Sevilla ⁵², L. Goossens ³², P.A. Gorbounov ⁹⁹,
 H.A. Gordon ²⁷, I. Gorelov ¹⁰⁷, B. Gorini ³², E. Gorini ^{76a,76b}, A. Gorišek ⁷⁸, A.T. Goshaw ⁴⁸, C. Gössling ⁴⁶,
 M.I. Gostkin ⁶⁸, C.R. Goudet ¹¹⁹, D. Goujdami ^{137c}, A.G. Goussiou ¹⁴⁰, N. Govender ^{147b,u}, E. Gozani ¹⁵⁴,
 L. Gruber ⁵⁷, I. Grabowska-Bold ^{41a}, P.O.J. Gradin ⁵⁸, J. Gramling ⁵², E. Gramstad ¹²¹, S. Grancagnolo ¹⁷,
 V. Gratchev ¹²⁵, P.M. Gravila ^{28f}, H.M. Gray ³², Z.D. Greenwood ^{82,v}, C. Grefe ²³, K. Gregersen ⁸¹,
 I.M. Gregor ⁴⁵, P. Grenier ¹⁴⁵, K. Grevtsov ⁵, J. Griffiths ⁸, A.A. Grillo ¹³⁹, K. Grimm ⁷⁵, S. Grinstein ^{13,w},
 Ph. Gris ³⁷, J.-F. Grivaz ¹¹⁹, S. Groh ⁸⁶, E. Gross ¹⁷⁵, J. Grosse-Knetter ⁵⁷, G.C. Grossi ⁸², Z.J. Grout ⁸¹,

- L. Guan ⁹², W. Guan ¹⁷⁶, J. Guenther ⁶⁵, F. Guescini ^{163a}, D. Guest ¹⁶⁶, O. Gueta ¹⁵⁵, B. Gui ¹¹³,
 E. Guido ^{53a,53b}, T. Guillemin ⁵, S. Guindon ², U. Gul ⁵⁶, C. Gumpert ³², J. Guo ^{36c}, W. Guo ⁹², Y. Guo ^{36a},
 R. Gupta ⁴³, S. Gupta ¹²², G. Gustavino ^{134a,134b}, P. Gutierrez ¹¹⁵, N.G. Gutierrez Ortiz ⁸¹, C. Gutschow ⁸¹,
 C. Guyot ¹³⁸, M.P. Guzik ^{41a}, C. Gwenlan ¹²², C.B. Gwilliam ⁷⁷, A. Haas ¹¹², C. Haber ¹⁶, H.K. Hadavand ⁸,
 A. Hadef ⁸⁸, S. Hageböck ²³, M. Hagihara ¹⁶⁴, H. Hakobyan ^{180,*}, M. Haleem ⁴⁵, J. Haley ¹¹⁶, G. Halladjian ⁹³,
 G.D. Hallewell ⁸⁸, K. Hamacher ¹⁷⁸, P. Hamal ¹¹⁷, K. Hamano ¹⁷², A. Hamilton ^{147a}, G.N. Hamity ¹⁴¹,
 P.G. Hamnett ⁴⁵, L. Han ^{36a}, S. Han ^{35a}, K. Hanagaki ^{69,x}, K. Hanawa ¹⁵⁷, M. Hance ¹³⁹, B. Haney ¹²⁴,
 P. Hanke ^{60a}, R. Hanna ¹³⁸, J.B. Hansen ³⁹, J.D. Hansen ³⁹, M.C. Hansen ²³, P.H. Hansen ³⁹, K. Hara ¹⁶⁴,
 A.S. Hard ¹⁷⁶, T. Harenberg ¹⁷⁸, F. Hariri ¹¹⁹, S. Harkusha ⁹⁵, R.D. Harrington ⁴⁹, P.F. Harrison ¹⁷³,
 N.M. Hartmann ¹⁰², M. Hasegawa ⁷⁰, Y. Hasegawa ¹⁴², A. Hasib ⁴⁹, S. Hassani ¹³⁸, S. Haug ¹⁸, R. Hauser ⁹³,
 L. Hauswald ⁴⁷, L.B. Havener ³⁸, M. Havranek ¹³⁰, C.M. Hawkes ¹⁹, R.J. Hawkings ³², D. Hayakawa ¹⁵⁹,
 D. Hayden ⁹³, C.P. Hays ¹²², J.M. Hays ⁷⁹, H.S. Hayward ⁷⁷, S.J. Haywood ¹³³, S.J. Head ¹⁹, T. Heck ⁸⁶,
 V. Hedberg ⁸⁴, L. Heelan ⁸, K.K. Heidegger ⁵¹, S. Heim ⁴⁵, T. Heim ¹⁶, B. Heinemann ^{45,y}, J.J. Heinrich ¹⁰²,
 L. Heinrich ¹¹², C. Heinz ⁵⁵, J. Hejbal ¹²⁹, L. Helary ³², A. Held ¹⁷¹, S. Hellman ^{148a,148b}, C. Helsens ³²,
 J. Henderson ¹²², R.C.W. Henderson ⁷⁵, Y. Heng ¹⁷⁶, S. Henkelmann ¹⁷¹, A.M. Henriques Correia ³²,
 S. Henrot-Versille ¹¹⁹, G.H. Herbert ¹⁷, H. Herde ²⁵, V. Herget ¹⁷⁷, Y. Hernández Jiménez ^{147c}, G. Herten ⁵¹,
 R. Hertenberger ¹⁰², L. Hervas ³², T.C. Herwig ¹²⁴, G.G. Hesketh ⁸¹, N.P. Hessey ^{163a}, J.W. Hetherly ⁴³,
 S. Higashino ⁶⁹, E. Higón-Rodríguez ¹⁷⁰, E. Hill ¹⁷², J.C. Hill ³⁰, K.H. Hiller ⁴⁵, S.J. Hillier ¹⁹, I. Hinchliffe ¹⁶,
 M. Hirose ⁵¹, D. Hirschbuehl ¹⁷⁸, B. Hiti ⁷⁸, O. Hladík ¹²⁹, X. Hoad ⁴⁹, J. Hobbs ¹⁵⁰, N. Hod ^{163a},
 M.C. Hodgkinson ¹⁴¹, P. Hodgson ¹⁴¹, A. Hoecker ³², M.R. Hoeferkamp ¹⁰⁷, F. Hoenig ¹⁰², D. Hohn ²³,
 T.R. Holmes ¹⁶, M. Homann ⁴⁶, S. Honda ¹⁶⁴, T. Honda ⁶⁹, T.M. Hong ¹²⁷, B.H. Hooberman ¹⁶⁹,
 W.H. Hopkins ¹¹⁸, Y. Horii ¹⁰⁵, A.J. Horton ¹⁴⁴, J-Y. Hostachy ⁵⁸, S. Hou ¹⁵³, A. Hoummada ^{137a},
 J. Howarth ⁴⁵, J. Hoya ⁷⁴, M. Hrabovsky ¹¹⁷, I. Hristova ¹⁷, J. Hrvnac ¹¹⁹, T. Hrynevich ⁹⁶,
 P.J. Hsu ⁶³, S.-C. Hsu ¹⁴⁰, Q. Hu ^{36a}, S. Hu ^{36c}, Y. Huang ^{35a}, Z. Hubacek ¹³⁰, F. Hubaut ⁸⁸, F. Huegging ²³,
 T.B. Huffman ¹²², E.W. Hughes ³⁸, G. Hughes ⁷⁵, M. Huhtinen ³², P. Huo ¹⁵⁰, N. Huseynov ^{68,b}, J. Huston ⁹³,
 J. Huth ⁵⁹, G. Iacobucci ⁵², G. Iakovidis ²⁷, I. Ibragimov ¹⁴³, L. Iconomidou-Fayard ¹¹⁹, P. Iengo ³²,
 O. Igonkina ^{109,z}, T. Iizawa ¹⁷⁴, Y. Ikegami ⁶⁹, M. Ikeno ⁶⁹, Y. Ilchenko ^{11,aa}, D. Iliadis ¹⁵⁶, N. Ilic ¹⁴⁵,
 G. Introzzi ^{123a,123b}, P. Ioannou ^{9,*}, M. Iodice ^{136a}, K. Iordanidou ³⁸, V. Ippolito ⁵⁹, N. Ishijima ¹²⁰,
 M. Ishino ¹⁵⁷, M. Ishitsuka ¹⁵⁹, C. Issever ¹²², S. Istin ^{20a}, F. Ito ¹⁶⁴, J.M. Iturbe Ponce ⁸⁷, R. Iuppa ^{162a,162b},
 H. Iwasaki ⁶⁹, J.M. Izen ⁴⁴, V. Izzo ^{106a}, S. Jabbar ³, P. Jackson ¹, V. Jain ², K.B. Jakobi ⁸⁶, K. Jakobs ⁵¹,
 S. Jakobsen ³², T. Jakoubek ¹²⁹, D.O. Jamin ¹¹⁶, D.K. Jana ⁸², R. Jansky ⁶⁵, J. Janssen ²³, M. Janus ⁵⁷,
 P.A. Janus ^{41a}, G. Jarlskog ⁸⁴, N. Javadov ^{68,b}, T. Javůrek ⁵¹, M. Javurkova ⁵¹, F. Jeanneau ¹³⁸, L. Jeanty ¹⁶,
 J. Jejelava ^{54a,ab}, A. Jelinskas ¹⁷³, P. Jenni ^{51,ac}, C. Jeske ¹⁷³, S. Jézéquel ⁵, H. Ji ¹⁷⁶, J. Jia ¹⁵⁰, H. Jiang ⁶⁷,
 Y. Jiang ^{36a}, Z. Jiang ¹⁴⁵, S. Jiggins ⁸¹, J. Jimenez Pena ¹⁷⁰, S. Jin ^{35a}, A. Jinaru ^{28b}, O. Jinnouchi ¹⁵⁹,
 H. Jivan ^{147c}, P. Johansson ¹⁴¹, K.A. Johns ⁷, C.A. Johnson ⁶⁴, W.J. Johnson ¹⁴⁰, K. Jon-And ^{148a,148b},
 R.W.L. Jones ⁷⁵, S. Jones ⁷, T.J. Jones ⁷⁷, J. Jongmanns ^{60a}, P.M. Jorge ^{128a,128b}, J. Jovicevic ^{163a}, X. Ju ¹⁷⁶,
 A. Juste Rozas ^{13,w}, M.K. Köhler ¹⁷⁵, A. Kaczmarska ⁴², M. Kado ¹¹⁹, H. Kagan ¹¹³, M. Kagan ¹⁴⁵,
 S.J. Kahn ⁸⁸, T. Kaji ¹⁷⁴, E. Kajomovitz ⁴⁸, C.W. Kalderon ⁸⁴, A. Kaluza ⁸⁶, S. Kama ⁴³, A. Kamenshchikov ¹³²,
 N. Kanaya ¹⁵⁷, S. Kaneti ³⁰, L. Kanjir ⁷⁸, V.A. Kantserov ¹⁰⁰, J. Kanzaki ⁶⁹, B. Kaplan ¹¹², L.S. Kaplan ¹⁷⁶,
 D. Kar ^{147c}, K. Karakostas ¹⁰, N. Karastathis ¹⁰, M.J. Kareem ⁵⁷, E. Karentzos ¹⁰, S.N. Karpov ⁶⁸,
 Z.M. Karpova ⁶⁸, K. Karthik ¹¹², V. Kartvelishvili ⁷⁵, A.N. Karyukhin ¹³², K. Kasahara ¹⁶⁴, L. Kashif ¹⁷⁶,
 R.D. Kass ¹¹³, A. Kastanas ¹⁴⁹, Y. Kataoka ¹⁵⁷, C. Kato ¹⁵⁷, A. Katre ⁵², J. Katzy ⁴⁵, K. Kawade ¹⁰⁵,
 K. Kawagoe ⁷³, T. Kawamoto ¹⁵⁷, G. Kawamura ⁵⁷, E.F. Kay ⁷⁷, V.F. Kazanin ^{111,c}, R. Keeler ¹⁷², R. Kehoe ⁴³,
 J.S. Keller ⁴⁵, J.J. Kempster ⁸⁰, H. Keoshkerian ¹⁶¹, O. Kepka ¹²⁹, B.P. Kerševan ⁷⁸, S. Kersten ¹⁷⁸,
 R.A. Keyes ⁹⁰, M. Khader ¹⁶⁹, F. Khalil-zada ¹², A. Khanov ¹¹⁶, A.G. Kharlamov ^{111,c}, T. Kharlamova ^{111,c},
 A. Khodinov ¹⁶⁰, T.J. Khoo ⁵², V. Khovanskiy ^{99,*}, E. Khramov ⁶⁸, J. Khubua ^{54b,ad}, S. Kido ⁷⁰, C.R. Kilby ⁸⁰,
 H.Y. Kim ⁸, S.H. Kim ¹⁶⁴, Y.K. Kim ³³, N. Kimura ¹⁵⁶, O.M. Kind ¹⁷, B.T. King ⁷⁷, D. Kirchmeier ⁴⁷, J. Kirk ¹³³,
 A.E. Kiryunin ¹⁰³, T. Kishimoto ¹⁵⁷, D. Kisielewska ^{41a}, K. Kiuchi ¹⁶⁴, O. Kivernyk ¹³⁸, E. Kladiva ^{146b},
 T. Klapdor-Kleingrothaus ⁵¹, M.H. Klein ³⁸, M. Klein ⁷⁷, U. Klein ⁷⁷, K. Kleinknecht ⁸⁶, P. Klimek ¹¹⁰,
 A. Klimentov ²⁷, R. Klingenberg ⁴⁶, T. Klioutchnikova ³², E.-E. Kluge ^{60a}, P. Kluit ¹⁰⁹, S. Kluth ¹⁰³,
 J. Knapik ⁴², E. Kneringer ⁶⁵, E.B.F.G. Knoops ⁸⁸, A. Knue ¹⁰³, A. Kobayashi ¹⁵⁷, D. Kobayashi ¹⁵⁹,
 T. Kobayashi ¹⁵⁷, M. Kobel ⁴⁷, M. Kocian ¹⁴⁵, P. Kodys ¹³¹, T. Koffman ³¹, E. Koffeman ¹⁰⁹, N.M. Köhler ¹⁰³,

- T. Koi 145, M. Kolb 60b, I. Koletsou 5, A.A. Komar 98,* , Y. Komori 157, T. Kondo 69, N. Kondrashova 36c, K. Köneke 51, A.C. König 108, T. Kono 69,ae , R. Konoplich 112,af , N. Konstantinidis 81, R. Kopeliansky 64, S. Koperny 41a, A.K. Kopp 51, K. Korcyl 42, K. Kordas 156, A. Korn 81, A.A. Korol 111,c , I. Korolkov 13, E.V. Korolkova 141, O. Kortner 103, S. Kortner 103, T. Kosek 131, V.V. Kostyukhin 23, A. Kotwal 48, A. Koulouris 10, A. Kourkoumeli-Charalampidi 123a,123b , C. Kourkoumelis 9, V. Kouskoura 27, A.B. Kowalewska 42, R. Kowalewski 172, T.Z. Kowalski 41a , C. Kozakai 157, W. Kozanecki 138, A.S. Kozhin 132, V.A. Kramarenko 101, G. Kramberger 78, D. Krasnopevtsev 100, M.W. Krasny 83, A. Krasznahorkay 32, D. Krauss 103, A. Kravchenko 27, J.A. Kremer 41a , M. Kretz 60c , J. Kretzschmar 77 , K. Kreutzfeldt 55, P. Krieger 161, K. Krizka 33, K. Kroeninger 46, H. Kroha 103, J. Kroll 124, J. Kroseberg 23, J. Krstic 14, U. Kruchonak 68, H. Krüger 23, N. Krumnack 67, M.C. Kruse 48, M. Kruskal 24, T. Kubota 91, H. Kucuk 81, S. Kuday 4b , J.T. Kuechler 178, S. Kuehn 51, A. Kugel 60c , F. Kuger 177, T. Kuhl 45, V. Kukhtin 68, R. Kukla 88, Y. Kulchitsky 95, S. Kuleshov 34b , Y.P. Kulinich 169, M. Kuna 134a,134b , T. Kunigo 71, A. Kupco 129, O. Kuprash 155, H. Kurashige 70, L.L. Kurchaninov 163a , Y.A. Kurochkin 95, M.G. Kurth 35a , V. Kus 129, E.S. Kuwertz 172, M. Kuze 159, J. Kvita 117, T. Kwan 172, D. Kyriazopoulos 141, A. La Rosa 103, J.L. La Rosa Navarro 26d , L. La Rotonda 40a,40b , C. Lacasta 170, F. Lacava 134a,134b , J. Lacey 45, H. Lacker 17, D. Lacour 83, E. Ladygin 68, R. Lafaye 5, B. Laforge 83, T. Lagouri 179, S. Lai 57, S. Lammers 64, W. Lampl 7, E. Lançon 27, U. Landgraf 51, M.P.J. Landon 79, M.C. Lanfermann 52, V.S. Lang 60a , J.C. Lange 13, A.J. Lankford 166, F. Lanni 27, K. Lantzsch 23, A. Lanza 123a , A. Lapertosa 53a,53b , S. Laplace 83, J.F. Laporte 138, T. Lari 94a , F. Lasagni Manghi 22a,22b , M. Lassnig 32, P. Laurelli 50, W. Lavrijsen 16, A.T. Law 139, P. Laycock 77, T. Lazovich 59, M. Lazzaroni 94a,94b , B. Le 91, O. Le Dortz 83, E. Le Guirriec 88, E.P. Le Quilleuc 138, M. LeBlanc 172, T. LeCompte 6, F. Ledroit-Guillon 58 , C.A. Lee 27, S.C. Lee 153, L. Lee 1, B. Lefebvre 90, G. Lefebvre 83, M. Lefebvre 172, F. Legger 102, C. Leggett 16, A. Lehan 77, G. Lehmann Miotto 32, X. Lei 7, W.A. Leight 45, A.G. Leister 179, M.A.L. Leite 26d , R. Leitner 131, D. Lellouch 175, B. Lemmer 57, K.J.C. Leney 81, T. Lenz 23, B. Lenzi 32, R. Leone 7, S. Leone 126a,126b , C. Leonidopoulos 49, G. Lerner 151, C. Leroy 97, A.A.J. Lesage 138, C.G. Lester 30, M. Levchenko 125, J. Levêque 5, D. Levin 92, L.J. Levinson 175, M. Levy 19, D. Lewis 79, B. Li 36a,s , Changqiao Li 36a , H. Li 150, L. Li 48, L. Li 36c , Q. Li 35a , S. Li 48, X. Li 36c , Y. Li 143, Z. Liang 35a , B. Liberti 135a , A. Liblong 161, K. Lie 169, J. Liebal 23, W. Liebig 15, A. Limosani 152, S.C. Lin 153,ag , T.H. Lin 86, B.E. Lindquist 150, A.E. Lionti 52, E. Lipeles 124, A. Lipniacka 15, M. Lisovyi 60b , T.M. Liss 169,ah , A. Lister 171, A.M. Litke 139, B. Liu 153,ai , H. Liu 92, H. Liu 27, J. Liu 36b , J.B. Liu 36a , K. Liu 88, L. Liu 169, M. Liu 36a , Y.L. Liu 36a , Y. Liu 36a , M. Livan 123a,123b , A. Lleres 58 , J. Llorente Merino 35a , S.L. Lloyd 79, C.Y. Lo 62b , F. Lo Sterzo 153, E.M. Lobodzinska 45, P. Loch 7, F.K. Loebinger 87, K.M. Loew 25, A. Loginov 179,* , T. Lohse 17, K. Lohwasser 45, M. Lokajicek 129, B.A. Long 24, J.D. Long 169, R.E. Long 75, L. Longo 76a,76b , K.A.Looper 113, J.A. Lopez 34b , D. Lopez Mateos 59, I. Lopez Paz 13, A. Lopez Solis 83, J. Lorenz 102, N. Lorenzo Martinez 64, M. Losada 21, P.J. Lösel 102, X. Lou 35a , A. Lounis 119, J. Love 6, P.A. Love 75, H. Lu 62a , N. Lu 92, Y.J. Lu 63, H.J. Lubatti 140, C. Luci 134a,134b , A. Lucotte 58 , C. Luedtke 51, F. Luehring 64, W. Lukas 65, L. Luminari 134a , O. Lundberg 148a,148b , B. Lund-Jensen 149, P.M. Luzi 83, D. Lynn 27, R. Lysak 129, E. Lytken 84, V. Lyubushkin 68, H. Ma 27, LL. Ma 36b , Y. Ma 36b , G. Maccarrone 50, A. Macchiolo 103, C.M. Macdonald 141, B. Maček 78, J. Machado Miguens 124,128b , D. Madaffari 88, R. Madar 37, H.J. Maddocks 168, W.F. Mader 47, A. Madsen 45, J. Maeda 70, S. Maeland 15, T. Maeno 27, A.S. Maevskiy 101, E. Magradze 57, J. Mahlstedt 109, C. Maiani 119, C. Maidantchik 26a , A.A. Maier 103, T. Maier 102, A. Maio 128a,128b,128d , S. Majewski 118, Y. Makida 69, N. Makovec 119, B. Malaescu 83, Pa. Malecki 42, V.P. Maleev 125, F. Malek 58 , U. Mallik 66, D. Malon 6, C. Malone 30, S. Maltezos 10, S. Malyukov 32, J. Mamuzic 170, G. Mancini 50, L. Mandelli 94a , I. Mandić 78 , J. Maneira 128a,128b , L. Manhaes de Andrade Filho 26b , J. Manjarres Ramos 163b , A. Mann 102, A. Manousos 32, B. Mansoulie 138, J.D. Mansour 35a , R. Mantifel 90, M. Mantoani 57, S. Manzoni 94a,94b , L. Mapelli 32, G. Marceca 29, L. March 52, G. Marchiori 83, M. Marcisovsky 129, M. Marjanovic 37, D.E. Marley 92, F. Marroquim 26a , S.P. Marsden 87, Z. Marshall 16, M.U.F. Martensson 168, S. Marti-Garcia 170, C.B. Martin 113, T.A. Martin 173, V.J. Martin 49, B. Martin dit Latour 15, M. Martinez 13,w , V.I. Martinez Outschoorn 169, S. Martin-Haugh 133, V.S. Martoiu 28b , A.C. Martyniuk 81, A. Marzin 115, L. Masetti 86, T. Mashimo 157, R. Mashinistov 98, J. Masik 87, A.L. Maslennikov 111,c , L. Massa 135a,135b , P. Mastrandrea 5, A. Mastroberardino 40a,40b , T. Masubuchi 157, P. Mättig 178, J. Maurer 28b , S.J. Maxfield 77, D.A. Maximov 111,c , R. Mazini 153, I. Maznas 156, S.M. Mazza 94a,94b , N.C. Mc Fadden 107, G. Mc Goldrick 161, S.P. Mc Kee 92, A. McCarn 92,

- R.L. McCarthy ¹⁵⁰, T.G. McCarthy ¹⁰³, L.I. McClymont ⁸¹, E.F. McDonald ⁹¹, J.A. McFayden ⁸¹,
 G. Mchedlidze ⁵⁷, S.J. McMahon ¹³³, P.C. McNamara ⁹¹, R.A. McPherson ^{172,o}, S. Meehan ¹⁴⁰, T.J. Megy ⁵¹,
 S. Mehlhase ¹⁰², A. Mehta ⁷⁷, T. Meideck ⁵⁸, K. Meier ^{60a}, C. Meineck ¹⁰², B. Meirose ⁴⁴, D. Melini ^{170,aj},
 B.R. Mellado Garcia ^{147c}, M. Melo ^{146a}, F. Meloni ¹⁸, S.B. Menary ⁸⁷, L. Meng ⁷⁷, X.T. Meng ⁹²,
 A. Mengarelli ^{22a,22b}, S. Menke ¹⁰³, E. Meoni ¹⁶⁵, S. Mergelmeyer ¹⁷, P. Mermod ⁵², L. Merola ^{106a,106b},
 C. Meroni ^{94a}, F.S. Merritt ³³, A. Messina ^{134a,134b}, J. Metcalfe ⁶, A.S. Mete ¹⁶⁶, C. Meyer ¹²⁴, J.-P. Meyer ¹³⁸,
 J. Meyer ¹⁰⁹, H. Meyer Zu Theenhausen ^{60a}, F. Miano ¹⁵¹, R.P. Middleton ¹³³, S. Miglioranzi ^{53a,53b},
 L. Mijović ⁴⁹, G. Mikenberg ¹⁷⁵, M. Mikestikova ¹²⁹, M. Mikuž ⁷⁸, M. Milesi ⁹¹, A. Milic ²⁷, D.W. Miller ³³,
 C. Mills ⁴⁹, A. Milov ¹⁷⁵, D.A. Milstead ^{148a,148b}, A.A. Minaenko ¹³², Y. Minami ¹⁵⁷, I.A. Minashvili ⁶⁸,
 A.I. Mincer ¹¹², B. Mindur ^{41a}, M. Mineev ⁶⁸, Y. Minegishi ¹⁵⁷, Y. Ming ¹⁷⁶, L.M. Mir ¹³, K.P. Mistry ¹²⁴,
 T. Mitani ¹⁷⁴, J. Mitrevski ¹⁰², V.A. Mitsou ¹⁷⁰, A. Miucci ¹⁸, P.S. Miyagawa ¹⁴¹, A. Mizukami ⁶⁹,
 J.U. Mjörnmark ⁸⁴, M. Mlynarikova ¹³¹, T. Moa ^{148a,148b}, K. Mochizuki ⁹⁷, P. Mogg ⁵¹, S. Mohapatra ³⁸,
 S. Molander ^{148a,148b}, R. Moles-Valls ²³, R. Monden ⁷¹, M.C. Mondragon ⁹³, K. Mönig ⁴⁵, J. Monk ³⁹,
 E. Monnier ⁸⁸, A. Montalbano ¹⁵⁰, J. Montejo Berlingen ³², F. Monticelli ⁷⁴, S. Monzani ^{94a,94b},
 R.W. Moore ³, N. Morange ¹¹⁹, D. Moreno ²¹, M. Moreno Llácer ⁵⁷, P. Morettini ^{53a}, S. Morgenstern ³²,
 D. Mori ¹⁴⁴, T. Mori ¹⁵⁷, M. Morii ⁵⁹, M. Morinaga ¹⁵⁷, V. Morisbak ¹²¹, A.K. Morley ¹⁵², G. Mornacchi ³²,
 J.D. Morris ⁷⁹, L. Morvaj ¹⁵⁰, P. Moschovakos ¹⁰, M. Mosidze ^{54b}, H.J. Moss ¹⁴¹, J. Moss ^{145,ak},
 K. Motohashi ¹⁵⁹, R. Mount ¹⁴⁵, E. Mountricha ²⁷, E.J.W. Moyse ⁸⁹, S. Muanza ⁸⁸, R.D. Mudd ¹⁹,
 F. Mueller ¹⁰³, J. Mueller ¹²⁷, R.S.P. Mueller ¹⁰², D. Muenstermann ⁷⁵, P. Mullen ⁵⁶, G.A. Mullier ¹⁸,
 F.J. Munoz Sanchez ⁸⁷, W.J. Murray ^{173,133}, H. Musheghyan ⁵⁷, M. Muškinja ⁷⁸, A.G. Myagkov ^{132,al},
 M. Myska ¹³⁰, B.P. Nachman ¹⁶, O. Nackenhorst ⁵², K. Nagai ¹²², R. Nagai ^{69,ae}, K. Nagano ⁶⁹, Y. Nagasaka ⁶¹,
 K. Nagata ¹⁶⁴, M. Nagel ⁵¹, E. Nagy ⁸⁸, A.M. Nairz ³², Y. Nakahama ¹⁰⁵, K. Nakamura ⁶⁹, T. Nakamura ¹⁵⁷,
 I. Nakano ¹¹⁴, R.F. Naranjo Garcia ⁴⁵, R. Narayan ¹¹, D.I. Narrias Villar ^{60a}, I. Naryshkin ¹²⁵, T. Naumann ⁴⁵,
 G. Navarro ²¹, R. Nayyar ⁷, H.A. Neal ⁹², P.Yu. Nechaeva ⁹⁸, T.J. Neep ¹³⁸, A. Negri ^{123a,123b}, M. Negrini ^{22a},
 S. Nektarijevic ¹⁰⁸, C. Nellist ¹¹⁹, A. Nelson ¹⁶⁶, S. Nemecsek ¹²⁹, P. Nemethy ¹¹², A.A. Nepomuceno ^{26a},
 M. Nessi ^{32,am}, M.S. Neubauer ¹⁶⁹, M. Neumann ¹⁷⁸, R.M. Neves ¹¹², P. Nevski ²⁷, P.R. Newman ¹⁹,
 T.Y. Ng ^{62c}, T. Nguyen Manh ⁹⁷, R.B. Nickerson ¹²², R. Nicolaïdou ¹³⁸, J. Nielsen ¹³⁹, V. Nikolaenko ^{132,al},
 I. Nikolic-Audit ⁸³, K. Nikolopoulos ¹⁹, J.K. Nilsen ¹²¹, P. Nilsson ²⁷, Y. Ninomiya ¹⁵⁷, A. Nisati ^{134a},
 N. Nishu ^{35c}, R. Nisius ¹⁰³, T. Nobe ¹⁵⁷, Y. Noguchi ⁷¹, M. Nomachi ¹²⁰, I. Nomidis ³¹, M.A. Nomura ²⁷,
 T. Nooney ⁷⁹, M. Nordberg ³², N. Norjoharuddeen ¹²², O. Novgorodova ⁴⁷, S. Nowak ¹⁰³, M. Nozaki ⁶⁹,
 L. Nozka ¹¹⁷, K. Ntekas ¹⁶⁶, E. Nurse ⁸¹, F. Nuti ⁹¹, D.C. O’Neil ¹⁴⁴, A.A. O’Rourke ⁴⁵, V. O’Shea ⁵⁶,
 F.G. Oakham ^{31,d}, H. Oberlack ¹⁰³, T. Obermann ²³, J. Ocariz ⁸³, A. Ochi ⁷⁰, I. Ochoa ³⁸, J.P. Ochoa-Ricoux ^{34a},
 S. Oda ⁷³, S. Odaka ⁶⁹, H. Ogren ⁶⁴, A. Oh ⁸⁷, S.H. Oh ⁴⁸, C.C. Ohm ¹⁶, H. Ohman ¹⁶⁸, H. Oide ^{53a,53b},
 H. Okawa ¹⁶⁴, Y. Okumura ¹⁵⁷, T. Okuyama ⁶⁹, A. Olariu ^{28b}, L.F. Oleiro Seabra ^{128a}, S.A. Olivares Pino ⁴⁹,
 D. Oliveira Damazio ²⁷, A. Olszewski ⁴², J. Olszowska ⁴², A. Onofre ^{128a,128e}, K. Onogi ¹⁰⁵, P.U.E. Onyisi ^{11,aa},
 M.J. Oreglia ³³, Y. Oren ¹⁵⁵, D. Orestano ^{136a,136b}, N. Orlando ^{62b}, R.S. Orr ¹⁶¹, B. Osculati ^{53a,53b,*},
 R. Ospanov ⁸⁷, G. Otero y Garzon ²⁹, H. Otono ⁷³, M. Ouchrif ^{137d}, F. Ould-Saada ¹²¹, A. Ouraou ¹³⁸,
 K.P. Oussoren ¹⁰⁹, Q. Ouyang ^{35a}, M. Owen ⁵⁶, R.E. Owen ¹⁹, V.E. Ozcan ^{20a}, N. Ozturk ⁸, K. Pachal ¹⁴⁴,
 A. Pacheco Pages ¹³, L. Pacheco Rodriguez ¹³⁸, C. Padilla Aranda ¹³, S. Pagan Griso ¹⁶, M. Paganini ¹⁷⁹,
 F. Paige ²⁷, P. Pais ⁸⁹, G. Palacino ⁶⁴, S. Palazzo ^{40a,40b}, S. Palestini ³², M. Palka ^{41b}, D. Pallin ³⁷,
 E. St. Panagiotopoulou ¹⁰, I. Panagoulias ¹⁰, C.E. Pandini ⁸³, J.G. Panduro Vazquez ⁸⁰, P. Pani ³²,
 S. Panitkin ²⁷, D. Pantea ^{28b}, L. Paolozzi ⁵², Th.D. Papadopoulou ¹⁰, K. Papageorgiou ^{9,t}, A. Paramonov ⁶,
 D. Paredes Hernandez ¹⁷⁹, A.J. Parker ⁷⁵, M.A. Parker ³⁰, K.A. Parker ⁴⁵, F. Parodi ^{53a,53b}, J.A. Parsons ³⁸,
 U. Parzefall ⁵¹, V.R. Pascuzzi ¹⁶¹, J.M. Pasner ¹³⁹, E. Pasqualucci ^{134a}, S. Passaggio ^{53a}, Fr. Pastore ⁸⁰,
 S. Patariaia ¹⁷⁸, J.R. Pater ⁸⁷, T. Pauly ³², J. Pearce ¹⁷², B. Pearson ¹⁰³, L.E. Pedersen ³⁹, S. Pedraza Lopez ¹⁷⁰,
 R. Pedro ^{128a,128b}, S.V. Peleganchuk ^{111,c}, O. Penc ¹²⁹, C. Peng ^{35a}, H. Peng ^{36a}, J. Penwell ⁶⁴, B.S. Peralva ^{26b},
 M.M. Perego ¹³⁸, D.V. Perepelitsa ²⁷, L. Perini ^{94a,94b}, H. Pernegger ³², S. Perrella ^{106a,106b}, R. Peschke ⁴⁵,
 V.D. Peshekhonov ^{68,*}, K. Peters ⁴⁵, R.F.Y. Peters ⁸⁷, B.A. Petersen ³², T.C. Petersen ³⁹, E. Petit ⁵⁸,
 A. Petridis ¹, C. Petridou ¹⁵⁶, P. Petroff ¹¹⁹, E. Petrolo ^{134a}, M. Petrov ¹²², F. Petrucci ^{136a,136b},
 N.E. Pettersson ⁸⁹, A. Peyaud ¹³⁸, R. Pezoa ^{34b}, P.W. Phillips ¹³³, G. Piacquadio ¹⁵⁰, E. Pianori ¹⁷³,
 A. Picazio ⁸⁹, E. Piccaro ⁷⁹, M.A. Pickering ¹²², R. Piegaia ²⁹, J.E. Pilcher ³³, A.D. Pilkington ⁸⁷, A.W.J. Pin ⁸⁷,
 M. Pinamonti ^{167a,167c,an}, J.L. Pinfold ³, H. Pirumov ⁴⁵, M. Pitt ¹⁷⁵, L. Plazak ^{146a}, M.-A. Pleier ²⁷,

- V. Pleskot ⁸⁶, E. Plotnikova ⁶⁸, D. Pluth ⁶⁷, P. Podberezko ¹¹¹, R. Poettgen ^{148a,148b}, L. Poggioli ¹¹⁹,
 D. Pohl ²³, G. Polesello ^{123a}, A. Poley ⁴⁵, A. Policicchio ^{40a,40b}, R. Polifka ³², A. Polini ^{22a}, C.S. Pollard ⁵⁶,
 V. Polychronakos ²⁷, K. Pommès ³², L. Pontecorvo ^{134a}, B.G. Pope ⁹³, G.A. Popeneciu ^{28d}, A. Poppleton ³²,
 S. Pospisil ¹³⁰, K. Potamianos ¹⁶, I.N. Potrap ⁶⁸, C.J. Potter ³⁰, G. Pouillard ³², J. Poveda ³²,
 M.E. Pozo Astigarraga ³², P. Pralavorio ⁸⁸, A. Pranko ¹⁶, S. Prell ⁶⁷, D. Price ⁸⁷, L.E. Price ⁶, M. Primavera ^{76a},
 S. Prince ⁹⁰, K. Prokofiev ^{62c}, F. Prokoshin ^{34b}, S. Protopopescu ²⁷, J. Proudfoot ⁶, M. Przybycien ^{41a},
 D. Puddu ^{136a,136b}, A. Puri ¹⁶⁹, P. Puzo ¹¹⁹, J. Qian ⁹², G. Qin ⁵⁶, Y. Qin ⁸⁷, A. Quadt ⁵⁷,
 M. Queitsch-Maitland ⁴⁵, D. Quilty ⁵⁶, S. Raddum ¹²¹, V. Radeka ²⁷, V. Radescu ¹²², S.K. Radhakrishnan ¹⁵⁰,
 P. Radloff ¹¹⁸, P. Rados ⁹¹, F. Ragusa ^{94a,94b}, G. Rahal ¹⁸¹, J.A. Raine ⁸⁷, S. Rajagopalan ²⁷,
 C. Rangel-Smith ¹⁶⁸, M.G. Ratti ^{94a,94b}, D.M. Rauch ⁴⁵, F. Rauscher ¹⁰², S. Rave ⁸⁶, T. Ravenscroft ⁵⁶,
 I. Ravinovich ¹⁷⁵, M. Raymond ³², A.L. Read ¹²¹, N.P. Readoff ⁷⁷, M. Reale ^{76a,76b}, D.M. Rebuzzi ^{123a,123b},
 A. Redelbach ¹⁷⁷, G. Redlinger ²⁷, R. Reece ¹³⁹, R.G. Reed ^{147c}, K. Reeves ⁴⁴, L. Rehnisch ¹⁷, J. Reichert ¹²⁴,
 A. Reiss ⁸⁶, C. Rembser ³², H. Ren ^{35a}, M. Rescigno ^{134a}, S. Resconi ^{94a}, E.D. Ressegueie ¹²⁴, S. Rettie ¹⁷¹,
 E. Reynolds ¹⁹, O.L. Rezanova ^{111,c}, P. Reznicek ¹³¹, R. Rezvani ⁹⁷, R. Richter ¹⁰³, S. Richter ⁸¹,
 E. Richter-Was ^{41b}, O. Ricken ²³, M. Ridel ⁸³, P. Rieck ¹⁰³, C.J. Riegel ¹⁷⁸, J. Rieger ⁵⁷, O. Rifki ¹¹⁵,
 M. Rijssenbeek ¹⁵⁰, A. Rimoldi ^{123a,123b}, M. Rimoldi ¹⁸, L. Rinaldi ^{22a}, B. Ristić ⁵², E. Ritsch ³², I. Riu ¹³,
 F. Rizatdinova ¹¹⁶, E. Rizvi ⁷⁹, C. Rizzi ¹³, R.T. Roberts ⁸⁷, S.H. Robertson ^{90,o}, A. Robichaud-Veronneau ⁹⁰,
 D. Robinson ³⁰, J.E.M. Robinson ⁴⁵, A. Robson ⁵⁶, C. Roda ^{126a,126b}, Y. Rodina ^{88,ao}, A. Rodriguez Perez ¹³,
 D. Rodriguez Rodriguez ¹⁷⁰, S. Roe ³², C.S. Rogan ⁵⁹, O. Røhne ¹²¹, J. Roloff ⁵⁹, A. Romanikou ¹⁰⁰,
 M. Romano ^{22a,22b}, S.M. Romano Saez ³⁷, E. Romero Adam ¹⁷⁰, N. Rompotis ⁷⁷, M. Ronzani ⁵¹, L. Roos ⁸³,
 S. Rosati ^{134a}, K. Rosbach ⁵¹, P. Rose ¹³⁹, N.-A. Rosien ⁵⁷, V. Rossetti ^{148a,148b}, E. Rossi ^{106a,106b},
 L.P. Rossi ^{53a}, J.H.N. Rosten ³⁰, R. Rosten ¹⁴⁰, M. Rotaru ^{28b}, I. Roth ¹⁷⁵, J. Rothberg ¹⁴⁰, D. Rousseau ¹¹⁹,
 A. Rozanov ⁸⁸, Y. Rozen ¹⁵⁴, X. Ruan ^{147c}, F. Rubbo ¹⁴⁵, F. Rühr ⁵¹, A. Ruiz-Martinez ³¹, Z. Rurikova ⁵¹,
 N.A. Rusakovich ⁶⁸, A. Ruschke ¹⁰², H.L. Russell ¹⁴⁰, J.P. Rutherford ⁷, N. Ruthmann ³², Y.F. Ryabov ¹²⁵,
 M. Rybar ¹⁶⁹, G. Rybkin ¹¹⁹, S. Ryu ⁶, A. Ryzhov ¹³², G.F. Rzechorz ⁵⁷, A.F. Saavedra ¹⁵², G. Sabato ¹⁰⁹,
 S. Sacerdoti ²⁹, H.F.-W. Sadrozinski ¹³⁹, R. Sadykov ⁶⁸, F. Safai Tehrani ^{134a}, P. Saha ¹¹⁰, M. Sahinsoy ^{60a},
 M. Saimpert ⁴⁵, T. Saito ¹⁵⁷, H. Sakamoto ¹⁵⁷, Y. Sakurai ¹⁷⁴, G. Salamanna ^{136a,136b}, J.E. Salazar Loyola ^{34b},
 D. Salek ¹⁰⁹, P.H. Sales De Bruin ¹⁴⁰, D. Salihagic ¹⁰³, A. Salnikov ¹⁴⁵, J. Salt ¹⁷⁰, D. Salvatore ^{40a,40b},
 F. Salvatore ¹⁵¹, A. Salvucci ^{62a,62b,62c}, A. Salzburger ³², D. Sammel ⁵¹, D. Sampsonidis ¹⁵⁶, J. Sánchez ¹⁷⁰,
 V. Sanchez Martinez ¹⁷⁰, A. Sanchez Pineda ^{106a,106b}, H. Sandaker ¹²¹, R.L. Sandbach ⁷⁹, C.O. Sander ⁴⁵,
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 R. Santonico ^{135a,135b}, H. Santos ^{128a}, I. Santoyo Castillo ¹⁵¹, K. Sapp ¹²⁷, A. Sapronov ⁶⁸,
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